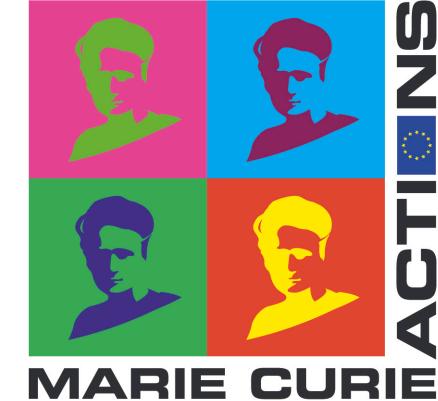


Jet quenching in evolving anisotropic matter

Andrey Sadofyev

IGFAE (USC)



Jet tomography

- Jets see the matter in HIC at multiple scales, and essentially X-ray it;
- The existing jet quenching theory is based on multiple simplifying assumptions: large parton energy, static matter, no fluctuations, etc;
- There is a very recent progress on the medium motion and structure effects in jet quenching and this is the focus of this talk;
- The developed formalism can be also applied to include orbital motion of nucleons and some of the in-medium fluctuations to the energy loss in cold nuclear matter;

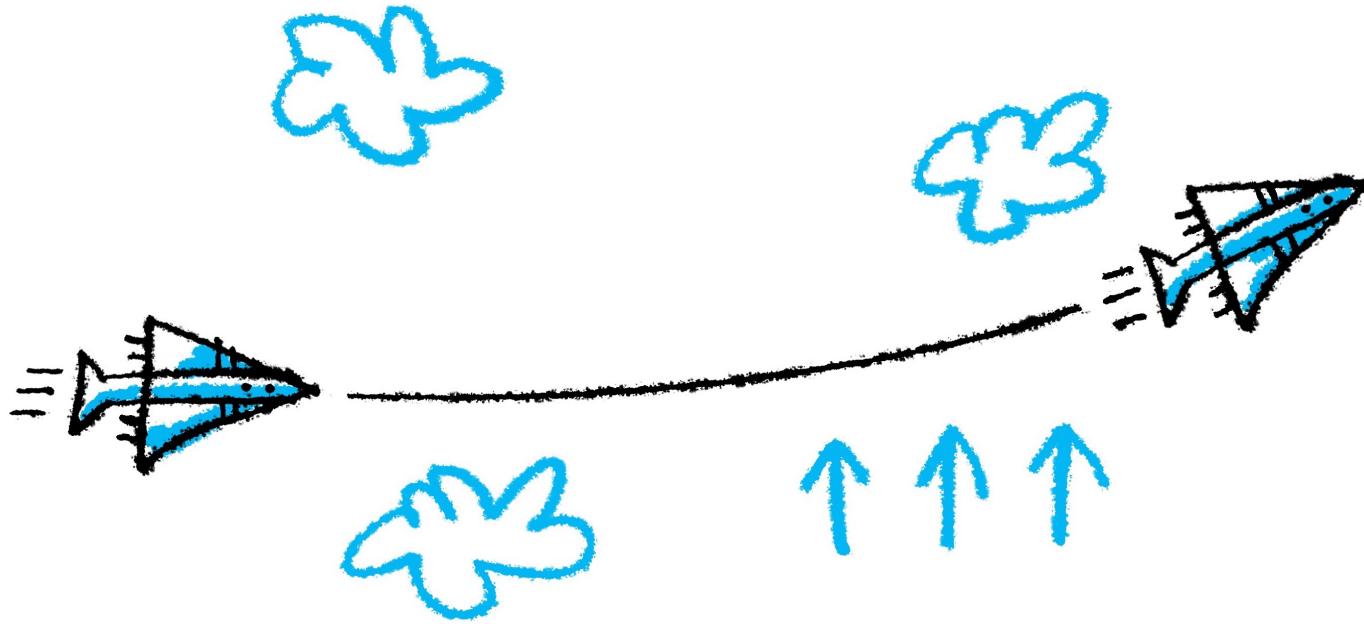




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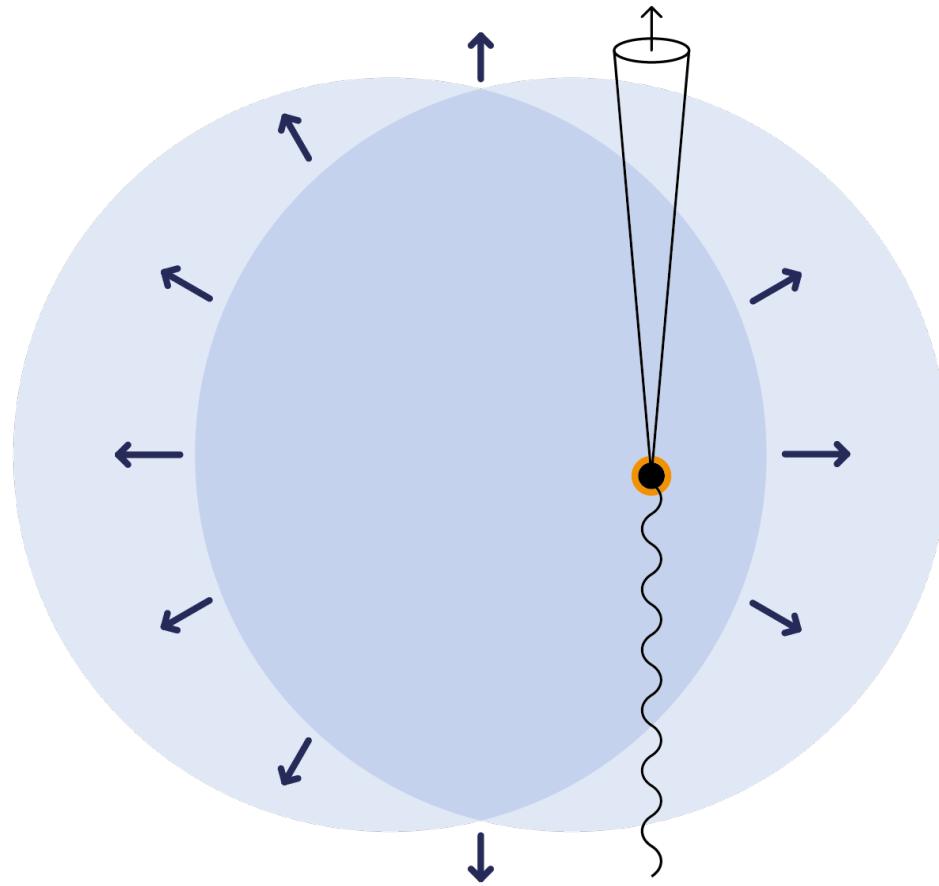
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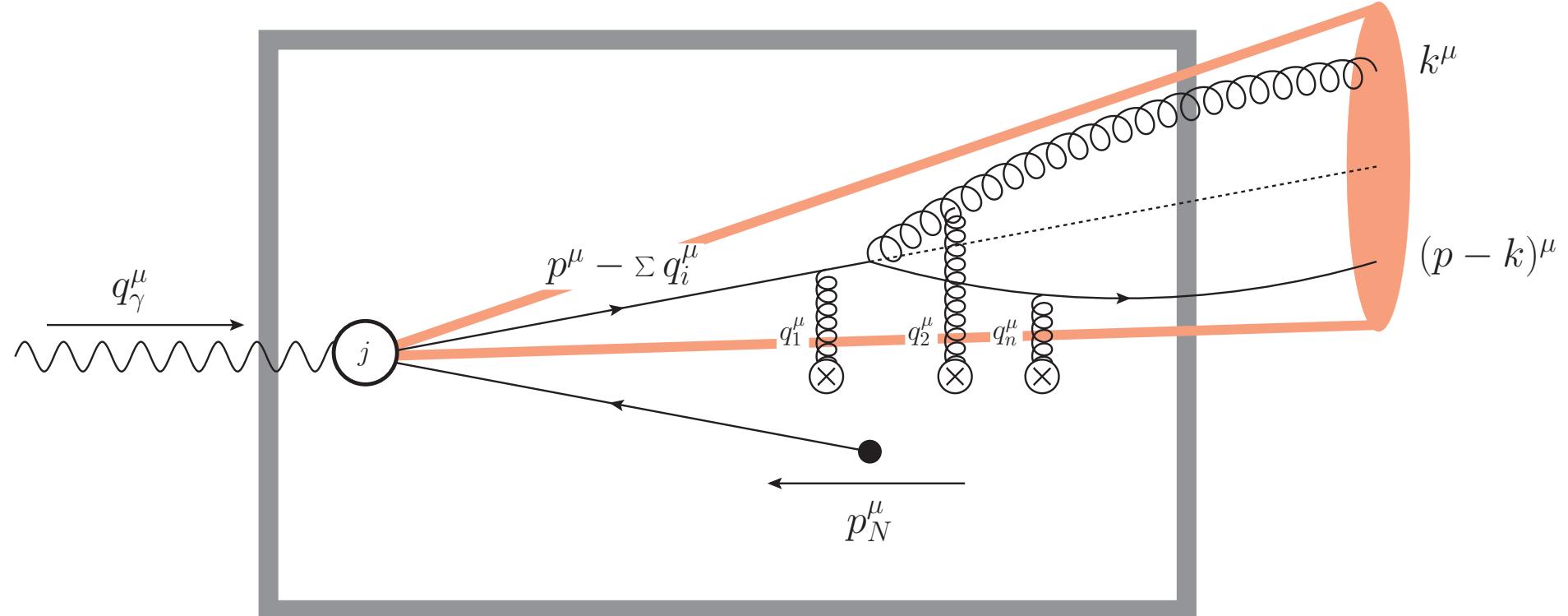


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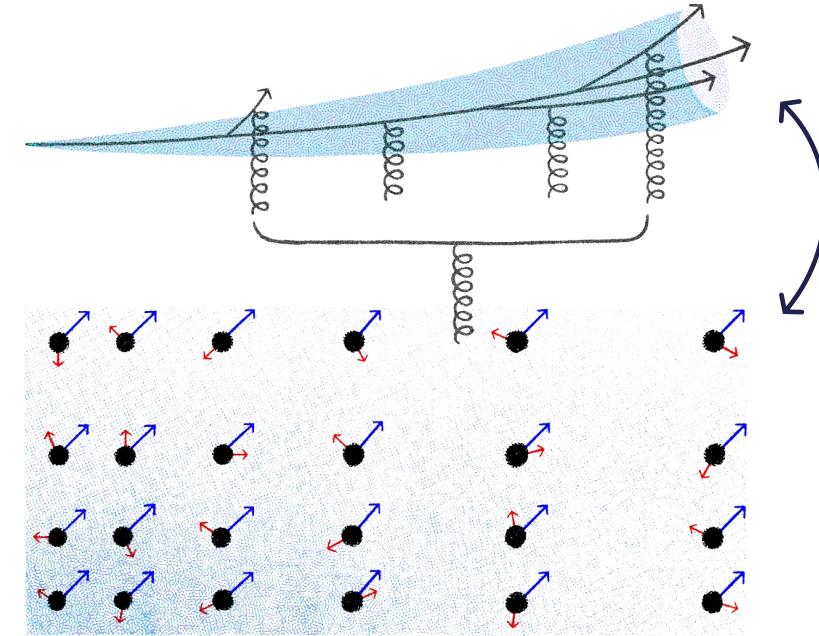
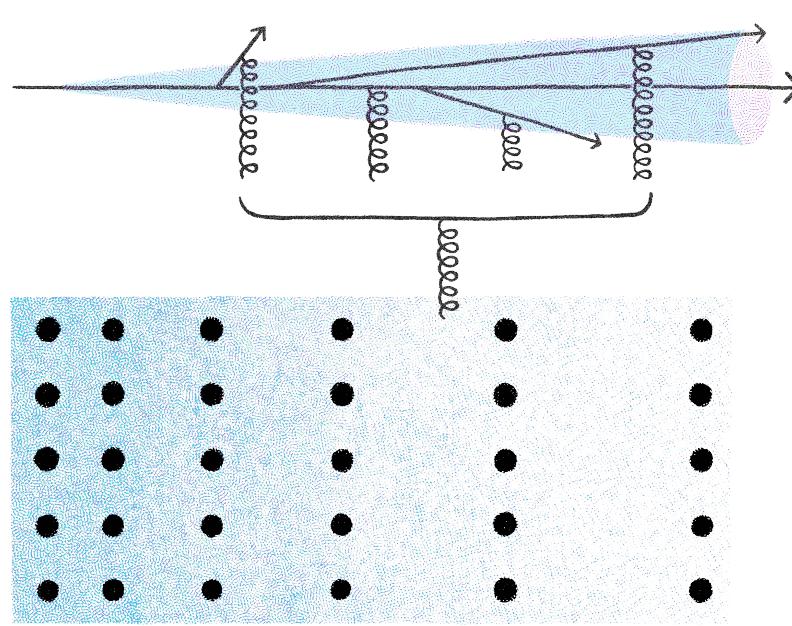




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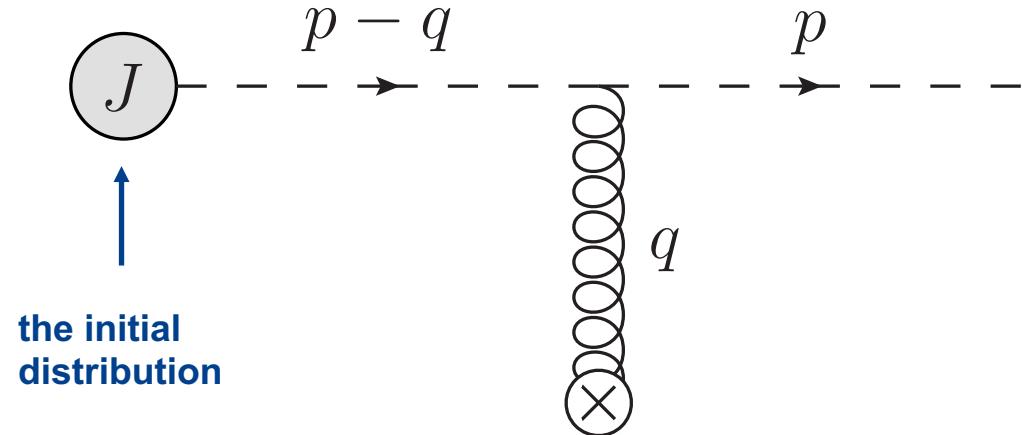
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Color potential

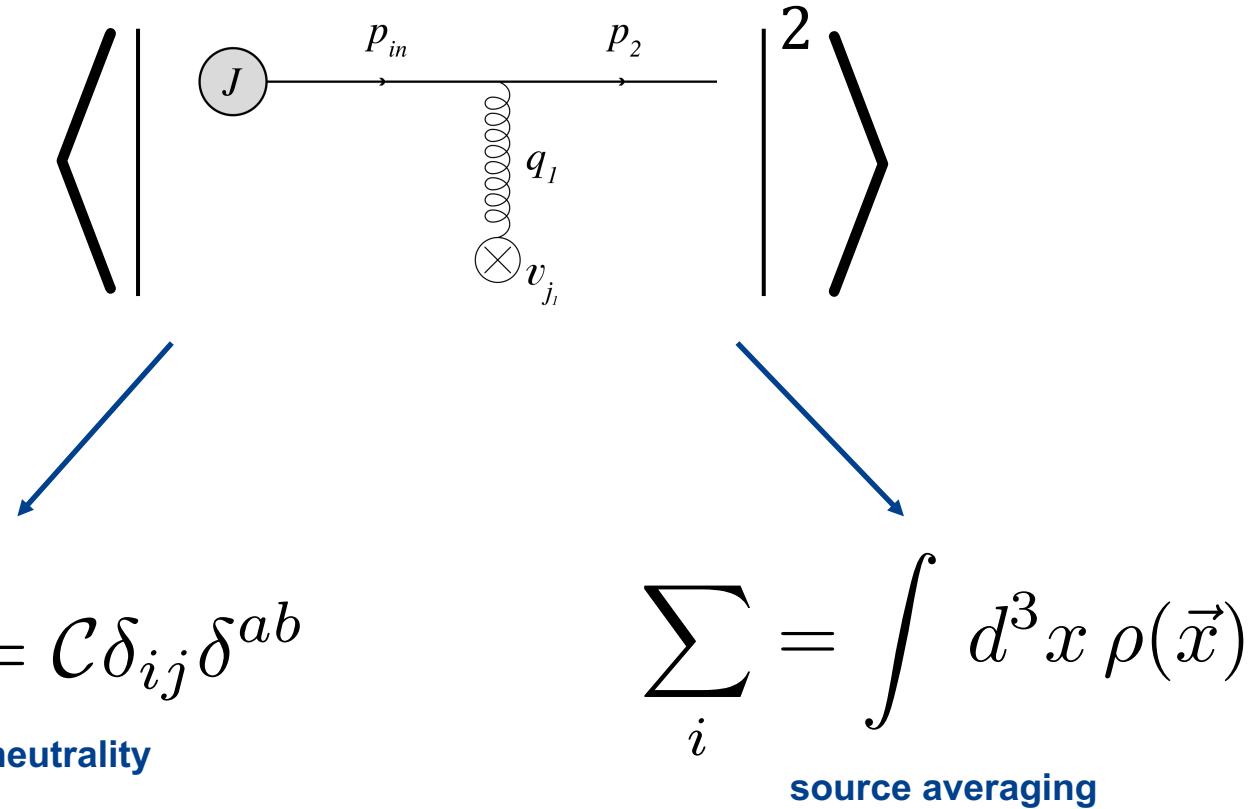


$$g A_{ext}^{\mu a}(q) = \sum_i e^{iq \cdot x_i} t_i^a u_i^\mu v_i(q) (2\pi) \delta(q^0 - \vec{u}_i \cdot \vec{q})$$

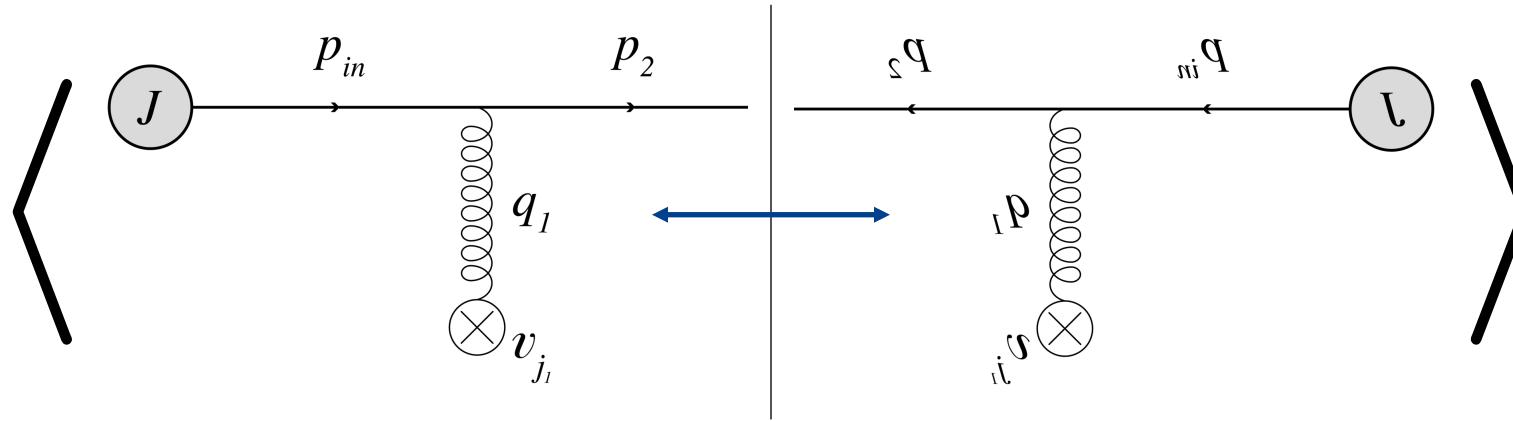
i μ a

inhomogeneity the fluid velocity

Medium averaging



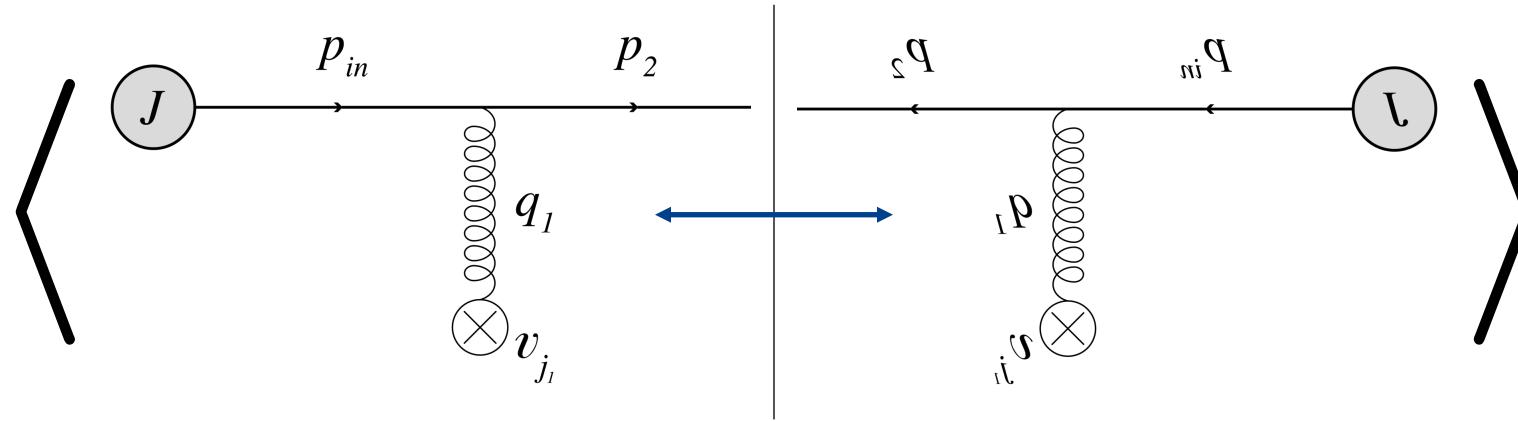
Medium averaging



$$\int d^2 \mathbf{x}_n e^{-i(\mathbf{q}_n \pm \bar{\mathbf{q}}_n) \cdot \mathbf{x}_n} = (2\pi)^2 \delta^{(2)}(\mathbf{q}_n \pm \bar{\mathbf{q}}_n)$$

$\rho \sim T^3$

Medium averaging

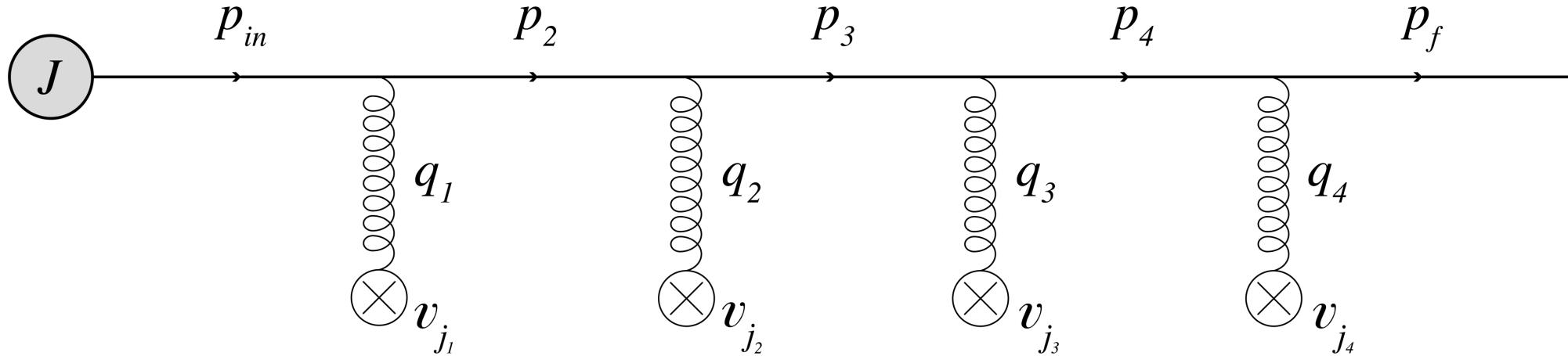


$$\int d^2 \mathbf{x}_n e^{-i(\mathbf{q}_n \pm \bar{\mathbf{q}}_n) \cdot \mathbf{x}_n} = (2\pi)^2 \delta^{(2)}(\mathbf{q}_n \pm \bar{\mathbf{q}}_n)$$

$$\int d^2 \mathbf{x}_n x_n^\alpha e^{-i(\mathbf{q}_n \pm \bar{\mathbf{q}}_n) \cdot \mathbf{x}_n} = i (2\pi)^2 \frac{\partial}{\partial(q_n \pm \bar{q}_n)_\alpha} \delta^{(2)}(\mathbf{q}_n \pm \bar{\mathbf{q}}_n)$$



The broadening

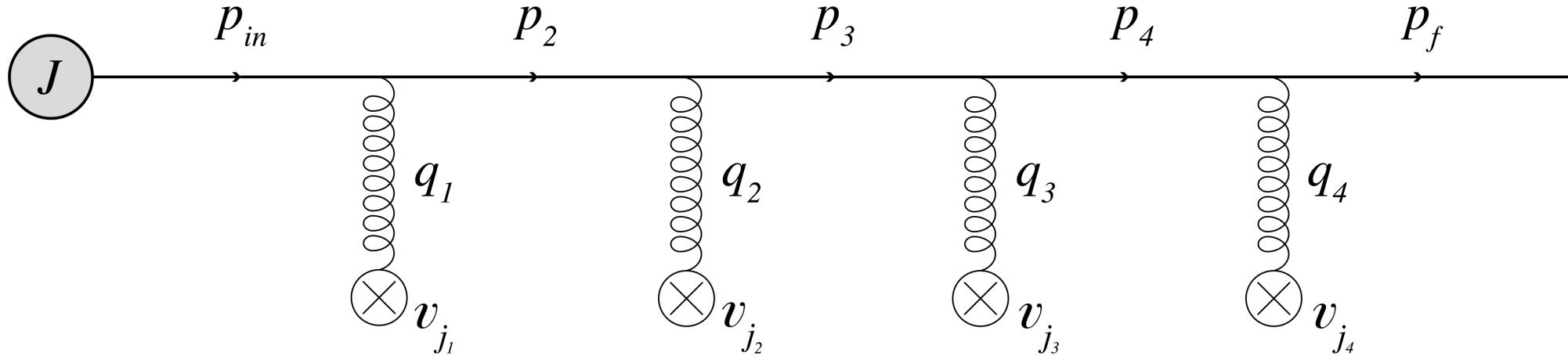


$$\begin{array}{c} E \rightarrow \infty \\ \downarrow \\ v(q) \\ \mu \ll E \quad \mu z \gg 1 \end{array} \longrightarrow iM(p) = \int \frac{d^2 \mathbf{p}_{in}}{(2\pi)^2} e^{i \frac{\mathbf{p}_f^2}{2E} L} G(\mathbf{p}_f, L; \mathbf{p}_{in}, 0) J(E, \mathbf{p}_{in})$$

↑
single particle propagator



The broadening



$$G(\mathbf{x}_L, L; \mathbf{x}_0, 0) = \int_{\mathbf{x}_0}^{\mathbf{x}_L} \mathcal{D}\mathbf{r} \exp \left(\frac{iE}{2} \int_0^L d\tau \dot{\mathbf{r}}^2 \right) \mathcal{P} \exp \left(-i \int_0^L d\tau t_{\text{proj}}^a v^a(\mathbf{r}(\tau), \tau) \right)$$



Jet broadening inhomogeneous matter

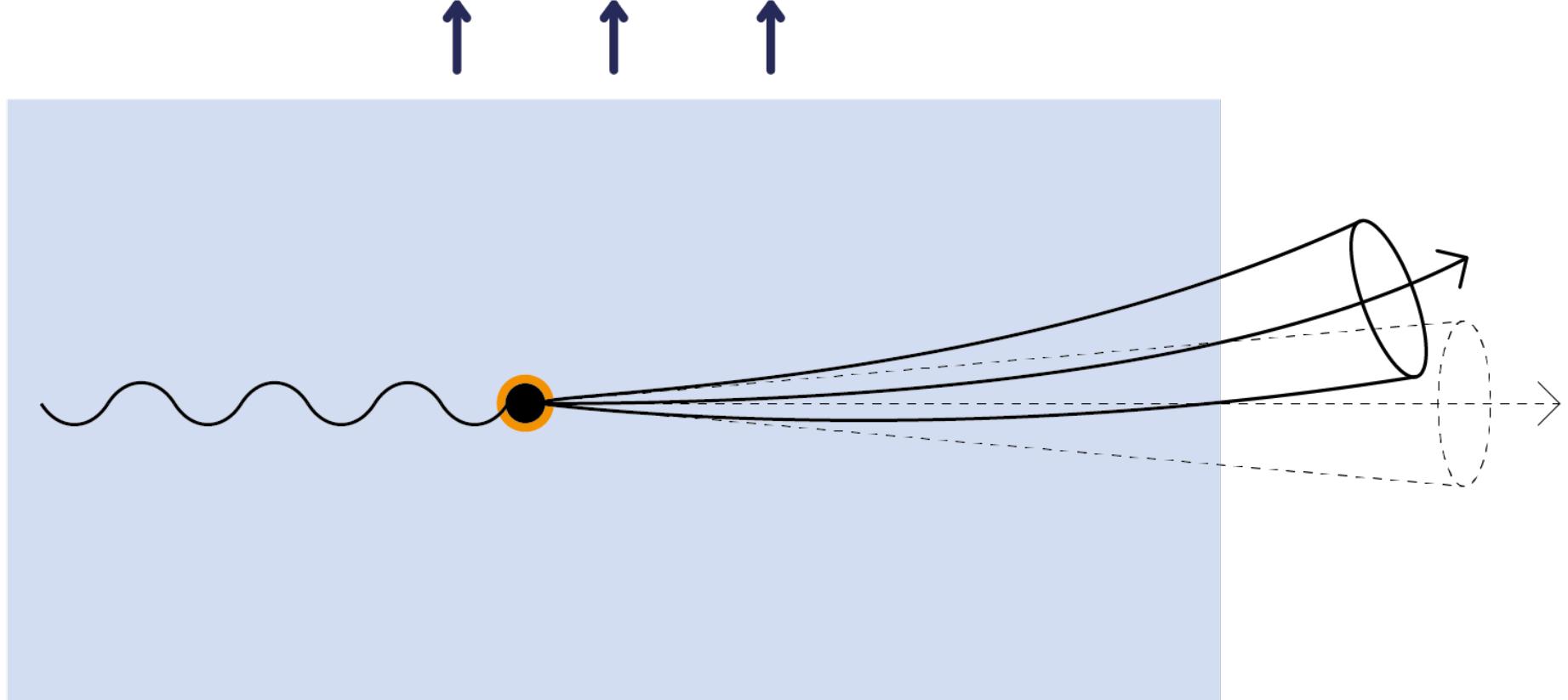
$$E \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}_0 dE} = \frac{1}{2(2\pi)^3} |J(p_0)|^2 = f(E) \delta^{(2)}(\mathbf{p}_0)$$

$$\frac{d\mathcal{N}}{d^2\mathbf{x} dE} \simeq \exp \{-\mathcal{V}(\mathbf{x}) L\} \left\{ \left[1 - \frac{iL^3}{6E} \nabla \mathcal{V}(\mathbf{x}) \cdot \hat{\mathbf{g}} \mathcal{V}(\mathbf{x}) \right] \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x} dE} + \frac{iL^2}{2E} \hat{\mathbf{g}} \mathcal{V}(\mathbf{x}) \cdot \nabla \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{x} dE} \right\}$$

$\hat{\mathbf{g}} \equiv \left(\nabla \rho \frac{\delta}{\delta \rho} + \nabla \mu^2 \frac{\delta}{\delta \mu^2} \right)$

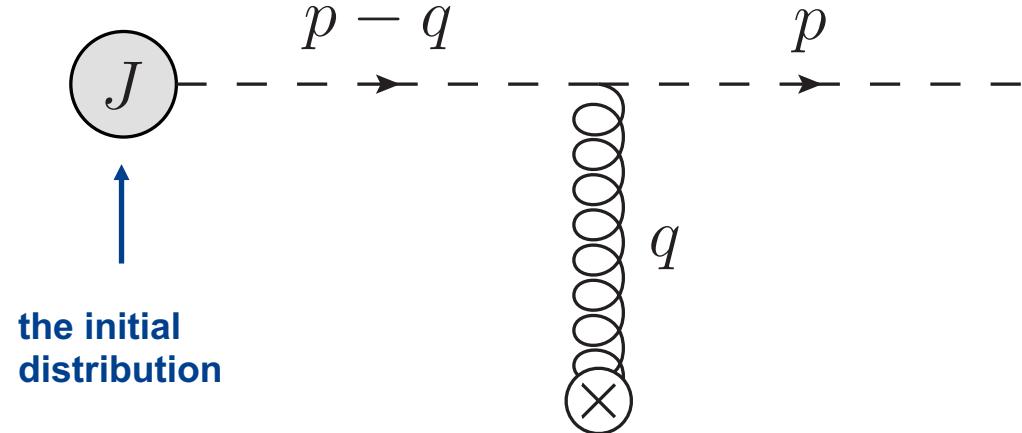
\downarrow
 $\rho \sim T^3$

$$\langle \mathbf{p} p_\perp^2 \rangle \simeq \chi^2 \frac{L \nabla T}{2T} \frac{\mu^4}{E} \left(\log \frac{E}{\mu} \right)^2$$



$$\langle \mathbf{p} p_{\perp}^2 \rangle \simeq \chi^2 \frac{L \nabla T}{2T} \frac{\mu^4}{E} \left(\log \frac{E}{\mu} \right)^2$$

Color potential



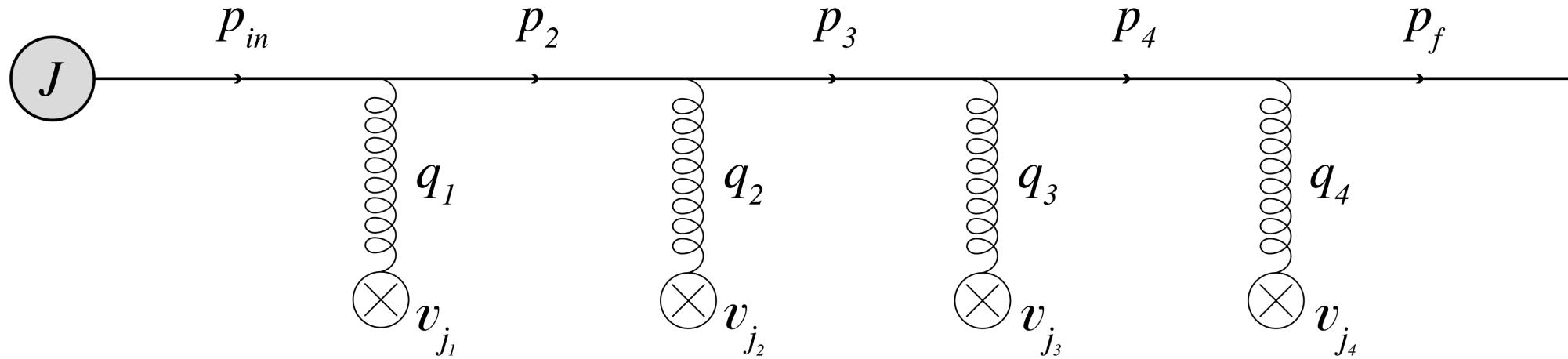
$$g A_{ext}^{\mu a}(q) = \sum_i e^{iq \cdot x_i} t_i^a u_i^\mu v_i(q) (2\pi) \delta(q^0 - \vec{u}_i \cdot \vec{q})$$

 **inhomogeneity**  **the fluid velocity**





Jet broadening



$$i\mathcal{M}_n(p) = e^{-i\left(\mathbf{u}\cdot\mathbf{p} - \frac{p_\perp^2}{2E}\right)L} \int \frac{d^2\mathbf{p}_0}{(2\pi)^2} \mathcal{G}_n(\mathbf{p}, L; \mathbf{p}_0, 0) J(E - \mathbf{u} \cdot (\mathbf{p} - \mathbf{p}_0), \mathbf{p}_0),$$



Jet broadening uniform matter

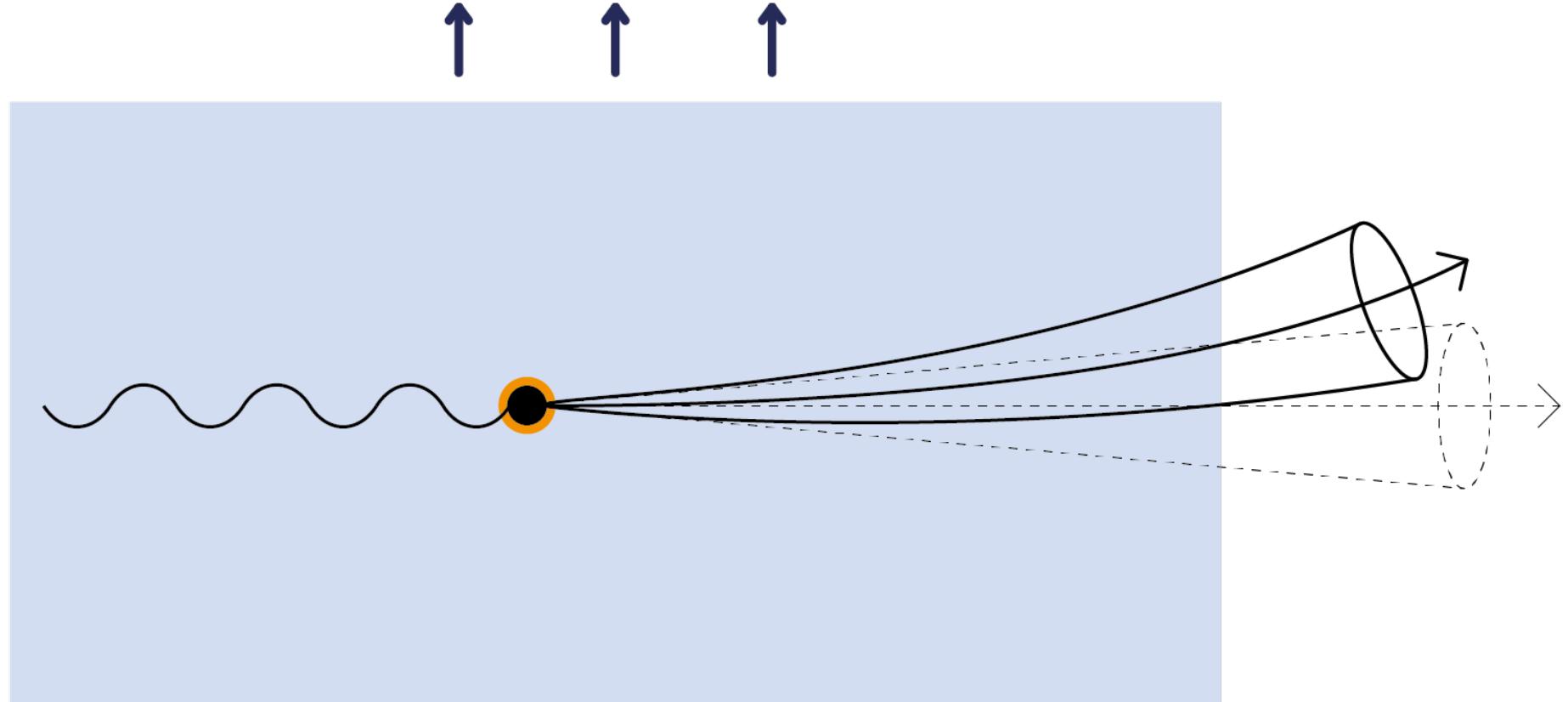
$$E \frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}_0 dE} = f(E) \delta^{(2)}(\mathbf{p}_0)$$

Eikonal approximation -- $E \rightarrow \infty$

$$\langle p_{\perp}^{2k} \mathbf{p} \rangle = \int \frac{d^2\mathbf{p} d^2\mathbf{r}}{(2\pi)^2} p_{\perp}^{2k} \mathbf{p} e^{-i\mathbf{p}\cdot\mathbf{r}} e^{-\mathcal{V}(\mathbf{r})L} = 0 + \mathcal{O}\left(\frac{\perp}{E}\right)$$

Opacity expansion -- $\chi \equiv \mathcal{C} \frac{g^4 \rho}{4\pi \mu^2} L \ll 1$

$$\langle p_{\perp}^{2k} \mathbf{p} \rangle \simeq -\frac{\mathbf{u}}{2E} \mathcal{C} \rho L \int \frac{d^2\mathbf{p}}{(2\pi)^2} p_{\perp}^{2k+2} \left[E \frac{f'(E)}{f(E)} v(p_{\perp})^2 + p_{\perp}^2 \frac{\partial v^2}{\partial p_{\perp}^2} \right]$$



$$\langle \mathbf{p} \rangle \simeq 3 \chi \mathbf{u} \frac{\mu^2}{E} \log \frac{E}{\mu}$$

Jet broadening

uniform matter

- Opacity $\chi \approx 4$
- $u \approx 0.7$ (about $\pi/4$ to z-axis)
- $\mu = gT$ with $g \approx 2$ and $T \approx 500 \text{ MeV}$

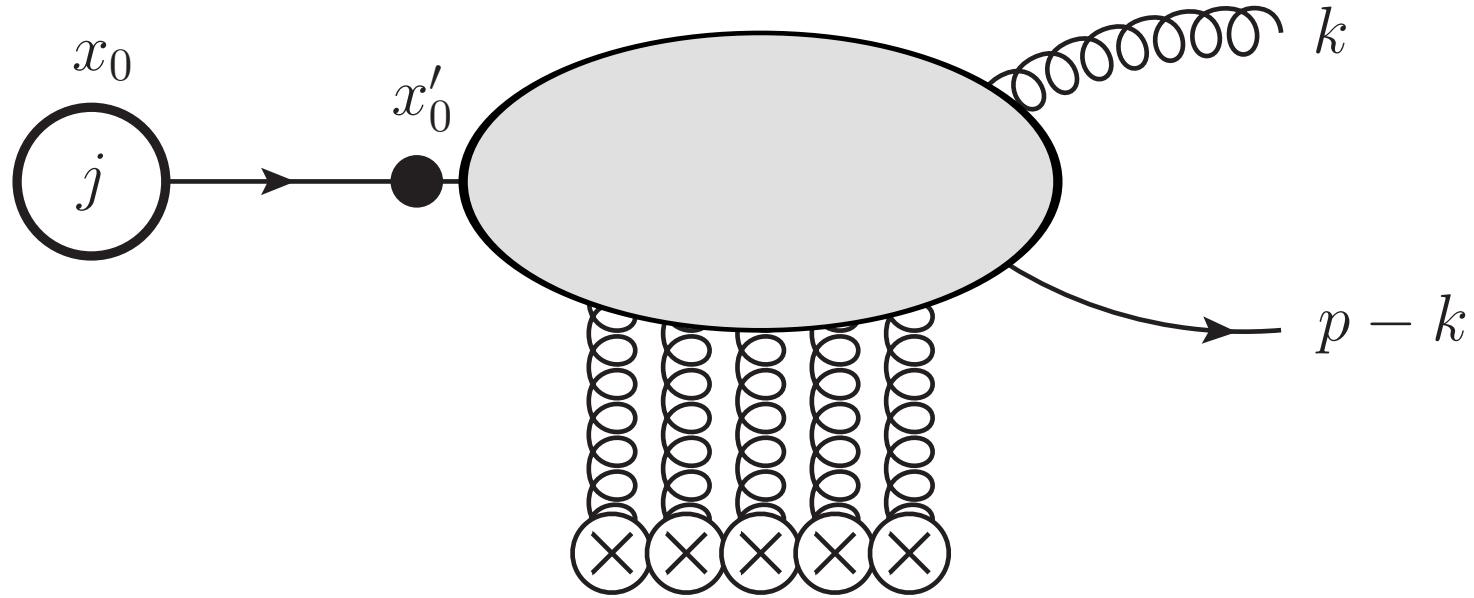
$$\left\langle \frac{p_\perp}{E} \right\rangle \simeq 3 \chi \frac{u_\perp}{1 - u_z} \frac{\mu^2}{E^2} \log \frac{E}{\mu}$$

What jet energy corresponds to $\langle \theta \rangle \approx 1^\circ$?



$E \sim 50 \text{ GeV}$

Gluon emission



Gluon emission

$$iR \simeq -\frac{g}{xE} \lim_{z_f \rightarrow \infty} \int_0^\infty dz_s \int d^2 x_0 e^{-i\mathbf{x}_0 \cdot \mathbf{l}_f} J(\mathbf{x}_0)$$

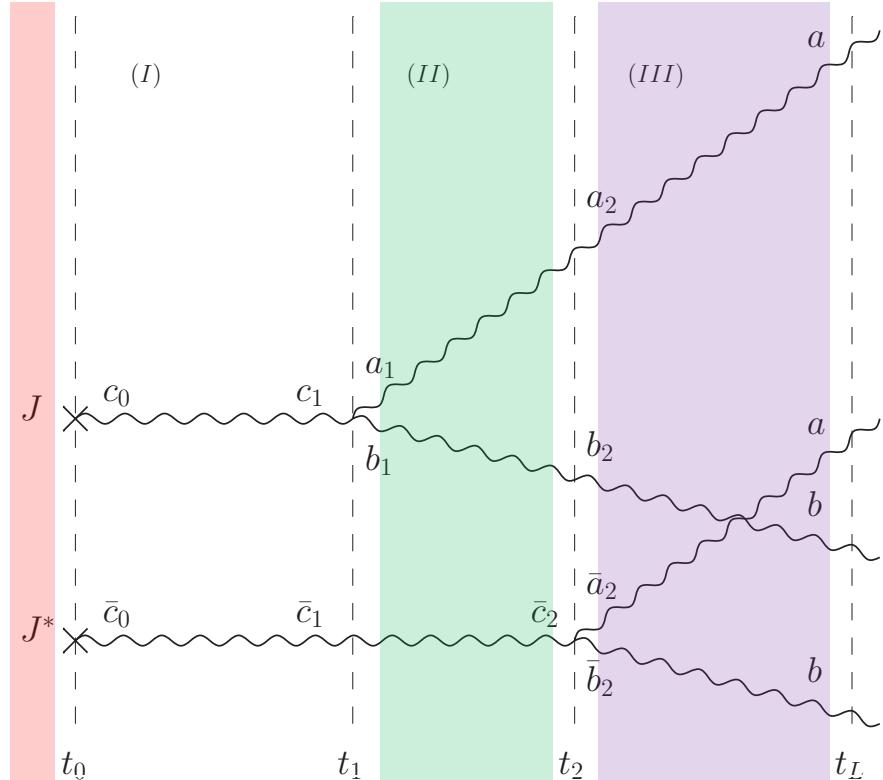
energy fraction  $\times \mathcal{W}(\mathbf{x}_0; \infty, z_s) t^a_{proj} \mathcal{W}(\mathbf{x}_0; z_s, 0) e^{i \frac{\mathbf{k}_f^2}{2\omega} z_f} \left[\epsilon \cdot \vec{\nabla}_{\mathbf{x}_0} \mathcal{G}^{ba} (\mathbf{k}_f, z_f; \mathbf{x}_0, z_s) \right]$



Wilson line ($x \ll 1$) the gluon (single-particle) propagator



Gluon emission



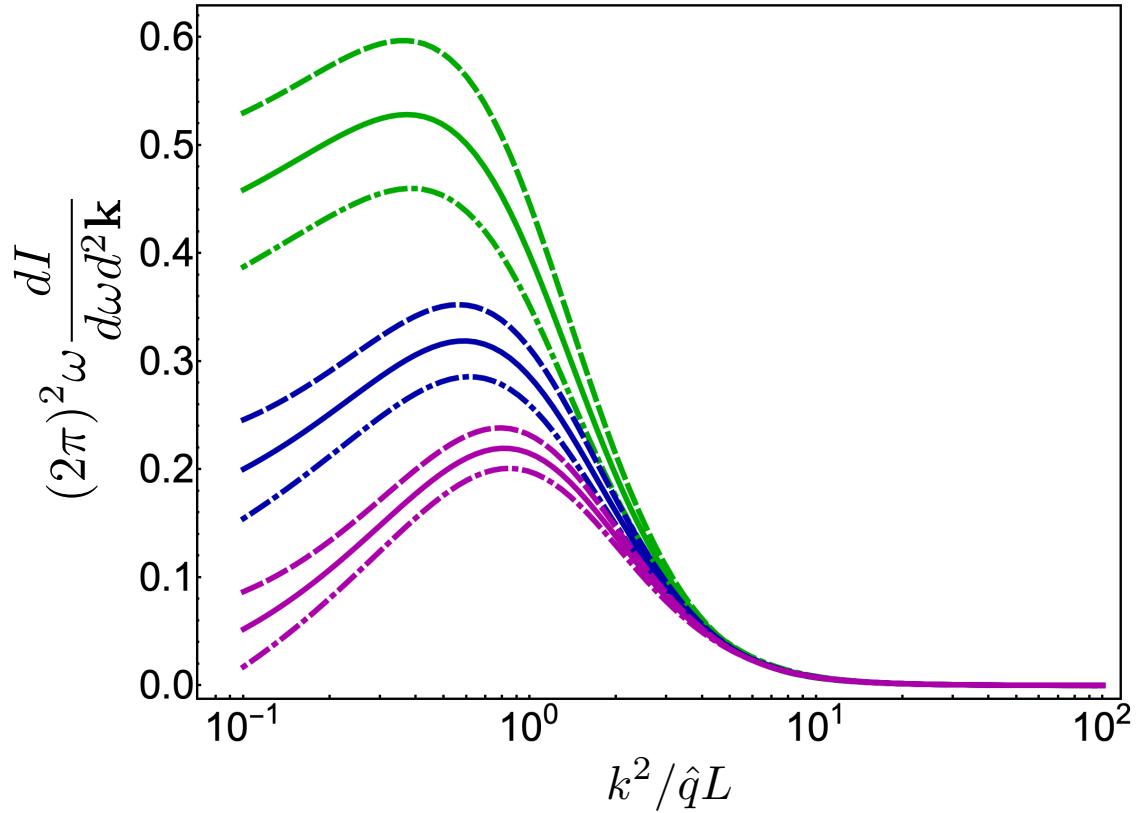
$$(2\pi)^2 \omega \frac{d\mathcal{N}}{d\omega d^2\mathbf{k}} = \frac{\alpha_s}{N_c \omega^2} \text{Re} \int_0^L d\bar{z} \int_0^{\bar{z}} dz \int_{\mathbf{x}_0, \mathbf{y}} |J(\mathbf{x}_0)|^2 \left[(\nabla_{\mathbf{x}} \cdot \nabla_{\bar{\mathbf{x}}}) \right. \\ \times \left. \left\langle \mathcal{G}^{bc}(\mathbf{k}, L; \mathbf{y}, \bar{z}) \mathcal{G}^{\dagger \bar{a}b}(\mathbf{k}, L; \bar{\mathbf{x}}, \bar{z}) \right\rangle \left\langle \mathcal{G}^{ca}(\mathbf{y}, \bar{z}; \mathbf{x}, z) \mathcal{W}_A^{\dagger a\bar{a}}(\mathbf{x}_0; \bar{z}, z) \right\rangle \right] \Big|_{\mathbf{x}=\bar{\mathbf{x}}=\mathbf{x}_0}$$

broadening of the gluon
(already have it with ∇)

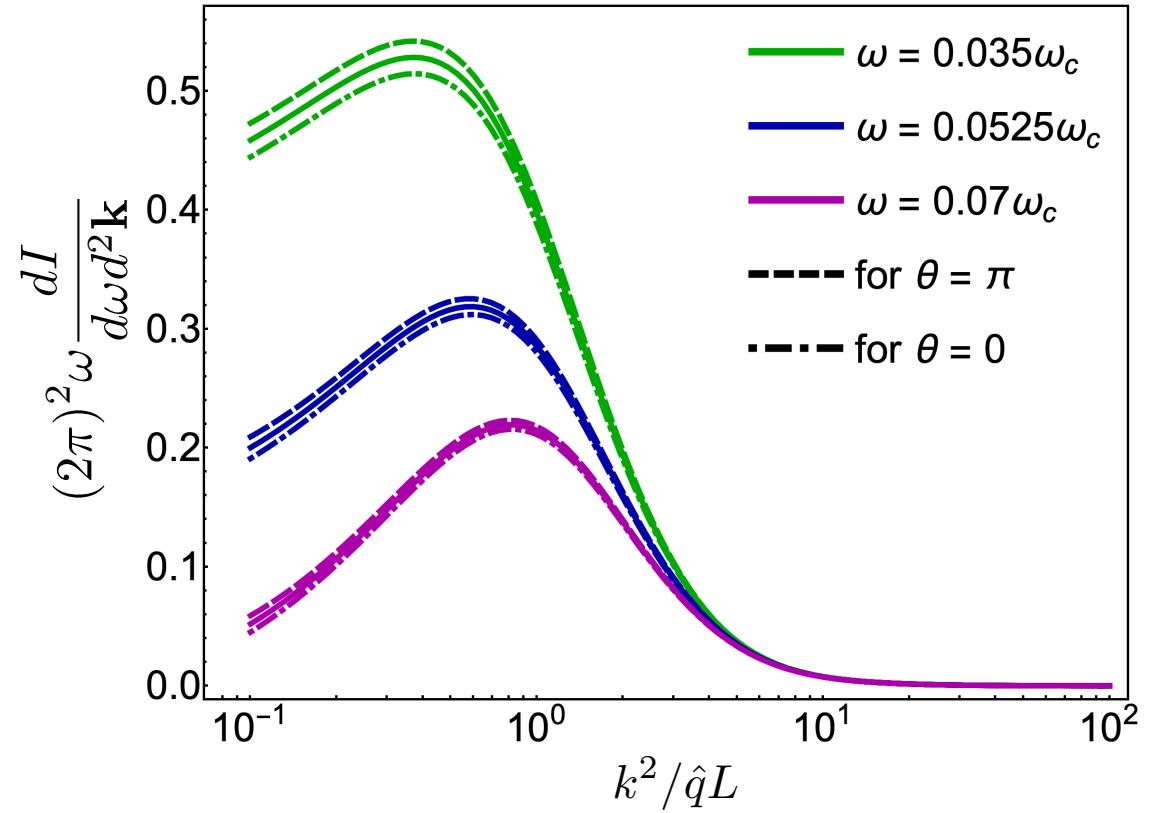
emission kernel

$$\delta\mathcal{K}(\mathbf{y}, \bar{z}; \mathbf{x}, z) = - \int_z^{\bar{z}} ds \int_{\mathbf{w}} \mathcal{K}_0(\mathbf{y}, \bar{z}; \mathbf{w}, s) \delta\mathcal{V}(\mathbf{w}, s) \mathcal{K}_0(\mathbf{w}, s; \mathbf{x}, z)$$

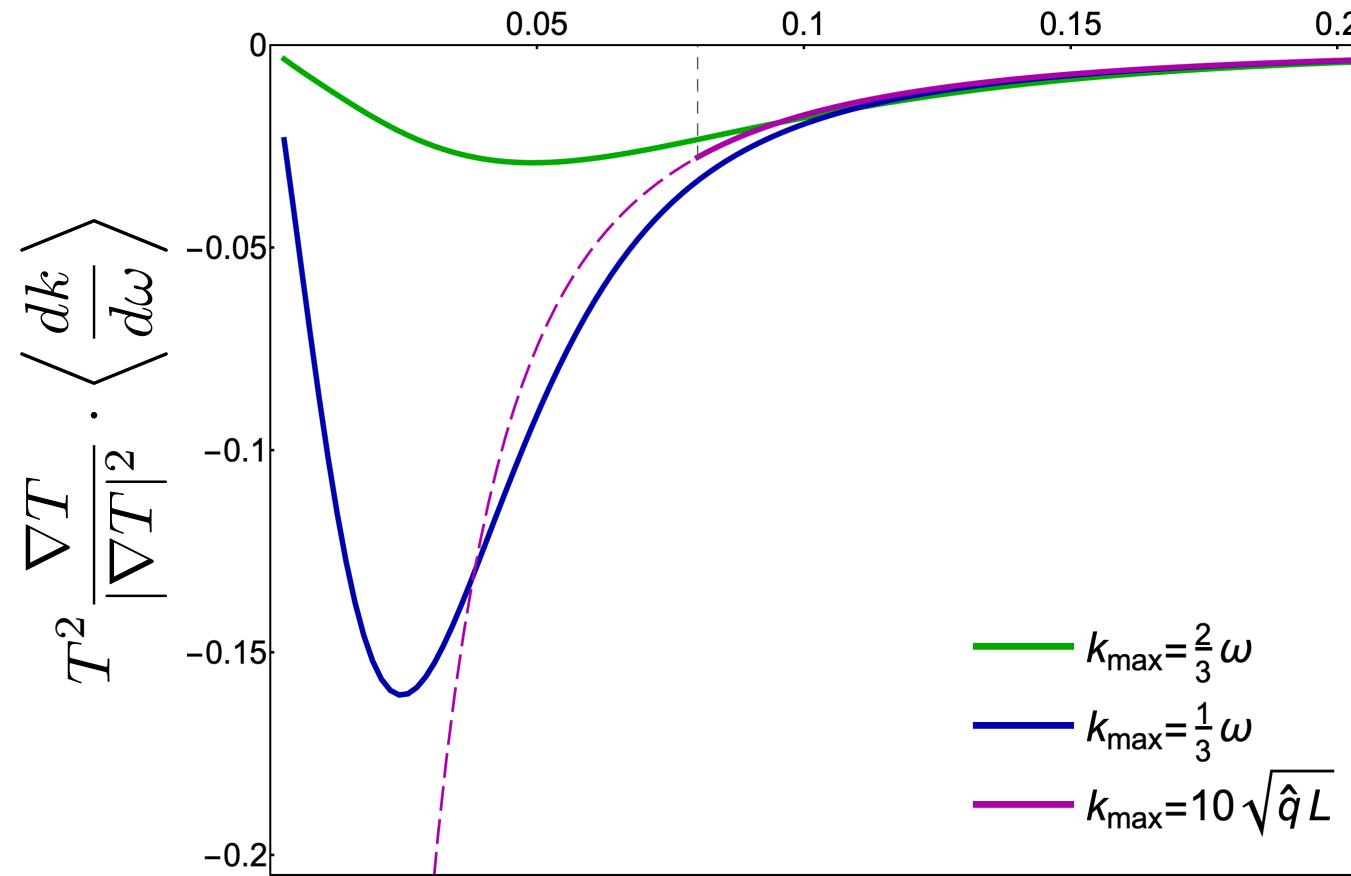
$$\left| \frac{\nabla T}{T^2} \right| = 0.05$$



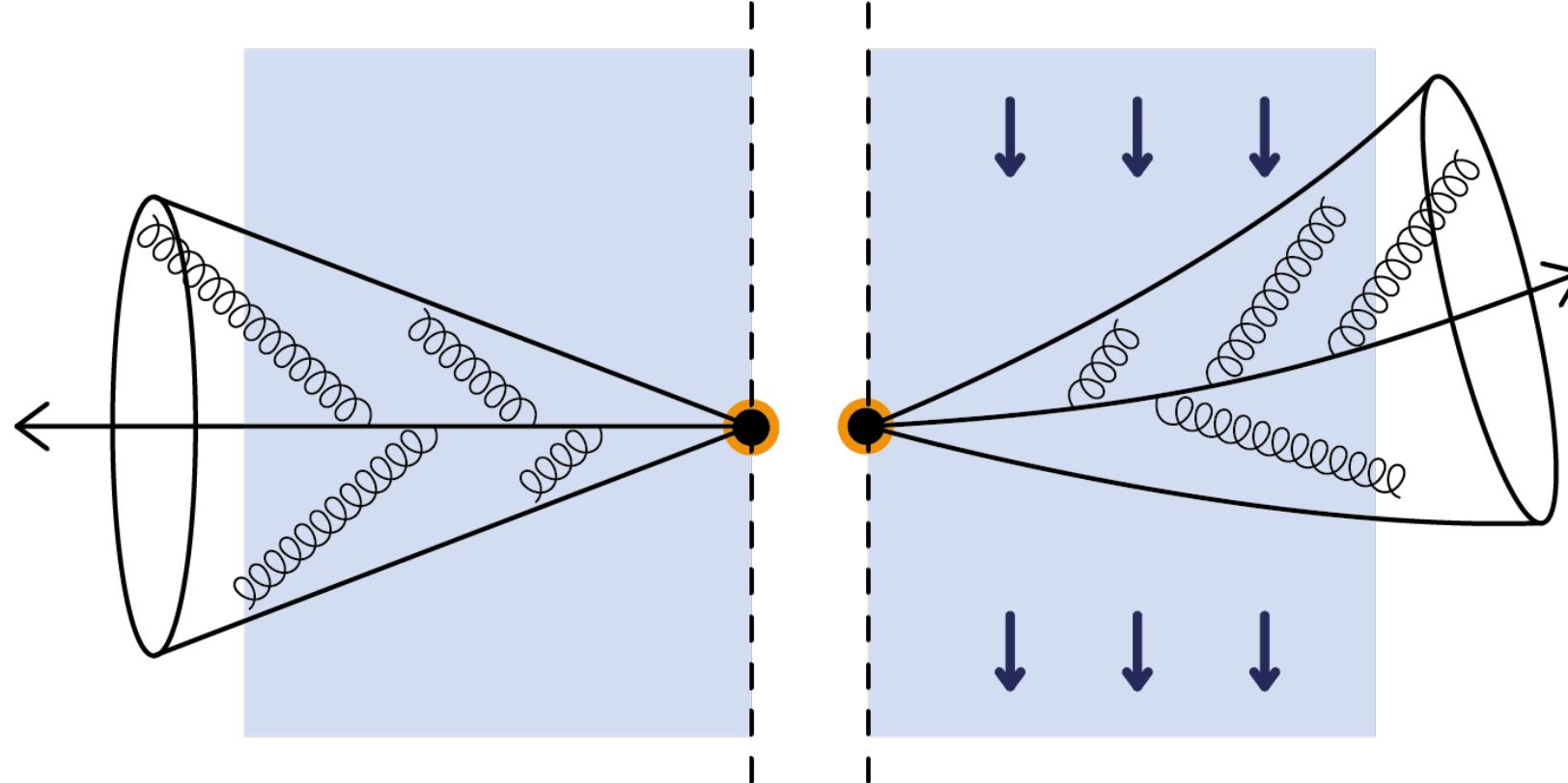
$$\left| \frac{\nabla T}{T^2} \right| = 0.01$$



$$L = 5\text{fm}, \quad T = 0.3\text{GeV}, \quad \hat{q} = 1\text{GeV}^2 \cdot \text{fm}^{-1}$$

 ω/ω_c 

$$\left\langle \frac{d\mathbf{k}}{d\omega} \right\rangle \equiv \int_{\Gamma} d^2\mathbf{k} \mathbf{k} \frac{dI}{d\omega d^2\mathbf{k}}$$



Summary

- Jets do feel the transverse flow and anisotropy, and get bended;
- The transverse flow and anisotropy do affect the medium-induced radiation, bending the substructure of jets;
- These effects can be in principle probed in experiment, leading us towards actual jet tomography;
- One should also expect similar evolution-induced effects for the other probes of nuclear matter;

