

# Direct-photon production and HBT correlations in Pb–Pb collisions at $\sqrt{s_{NN}} = 5.02$ (2.76) TeV with the ALICE experiment

A. Marin for the ALICE Collaboration



# Motivation: Direct photons

Carry information on the medium's temperature and space-time evolution.

Large background from neutral meson decays ( $\pi^0, \eta, \omega, \dots$ ).

## Prompt photons: [F. Jonas, talk 74](#)

- Dominant at high  $p_T$  ( $p_T > \sim 5$  GeV/c), power-law shape
- Initial hard scattering
- Described by NLO pQCD

## Pre-equilibrium photons:

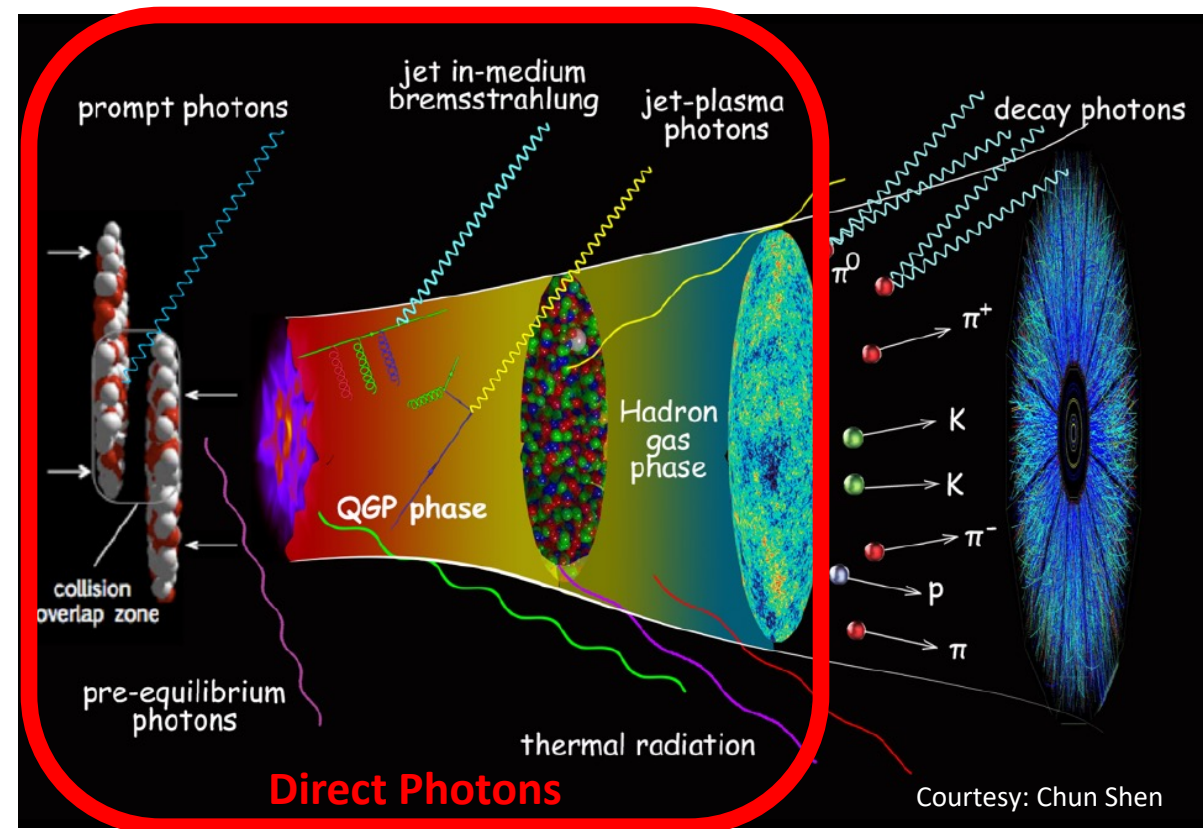
- Sensitive to the saturation momentum

## Jet-medium interactions:

- Scattering of hard partons with thermalized partons

## Thermal photons:

- Dominant at low  $p_T$  ( $p_T < \sim 3$  GeV/c) with exponential shape
- Emitted by thermalized medium
- Comparison to models employing hydrodynamics



# Measurement of inclusive photons

## EMCal/DCal: sampling calorimeter

10 modules at 4.4 m from ALICE IP.

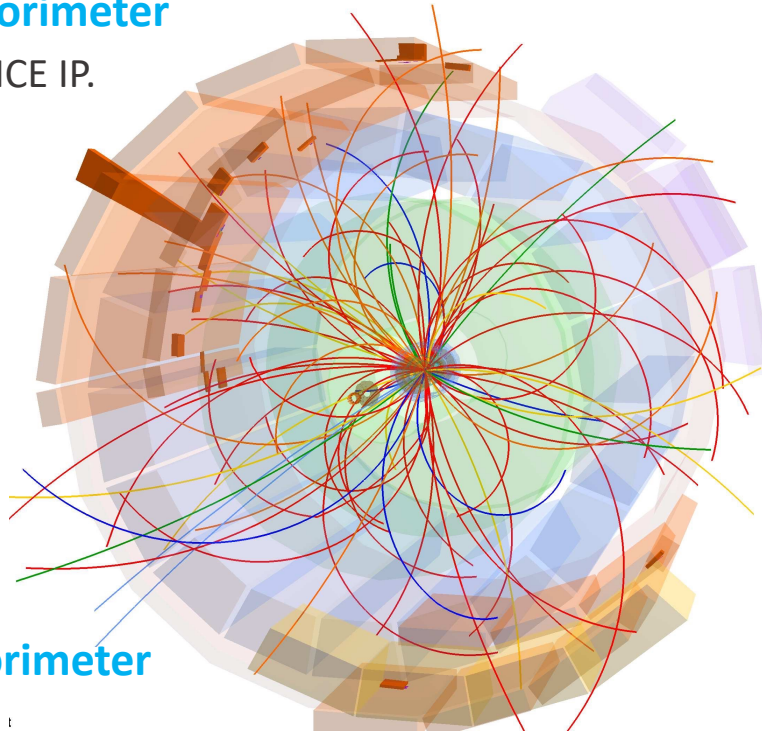
EMCal:

$|\eta| < 0.7, 80^\circ < \varphi < 187^\circ$ .

DCAL:

$0.22 < |\eta| < 0.7, 260^\circ < \varphi < 320^\circ$

$|\eta| < 0.7, 320^\circ < \varphi < 327^\circ$



## PHOS: homogeneous calorimeter

PbWO<sub>4</sub> crystal

3 modules at 4.6 m from ALICE IP

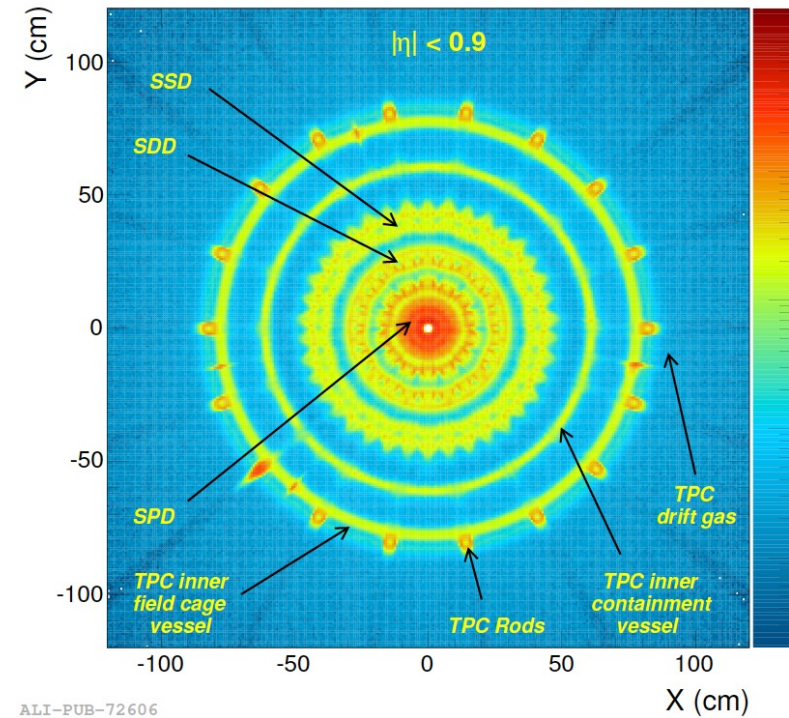
$|\eta| < 0.12, 260^\circ < \varphi < 320^\circ$

[J. Lühder, poster 52](#)

[N. Stangmann, poster 105](#)

[J. Koenig, poster 109](#)

$$R_\gamma = \frac{\gamma_{inc}}{\pi^0} / \frac{\gamma_{decay}}{\pi^0_{param}} \sim \frac{\gamma_{inc}}{\gamma_{decay}}$$



ALI-PUB-72606

## Photon conversion method (PCM):

Photon conversion in detector material ITS and TPC

$|\eta| < 0.9, R < 180 \text{ cm}, 0^\circ < \varphi < 360^\circ, X/X_0 = 11.4 \pm 0.5 \text{ sys } \%$

**10-15% low  $p_T$  direct photon excess at LHC energies**

**6% uncertainty, largest contribution: 4.5% sys  $X/X_0$**

Can this uncertainty be reduced? → Improve  $R_\gamma$  uncertainty

# Data-driven precision determination of the material budget in ALICE



RD: Real Data  
MC: Monte Carlo

[arxiv: 2303.15317](https://arxiv.org/abs/2303.15317)

$\omega_i$  : TPC-gas based calibration weights

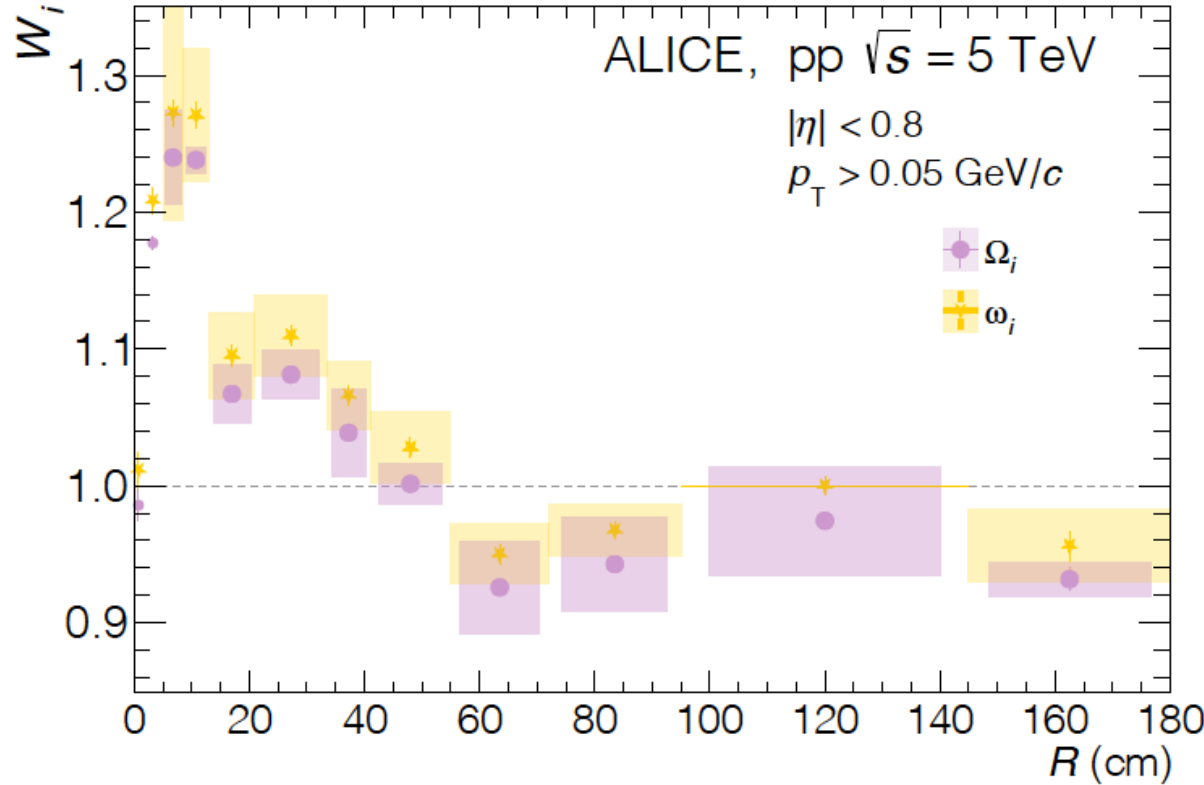
$$\omega_i = \frac{N_{\gamma,i}^{\text{rec,RD}} / N_{\gamma,\text{gas}}^{\text{rec,RD}}}{N_{\gamma,i}^{\text{rec,MC}} / N_{\gamma,\text{gas}}^{\text{rec,MC}}}$$

$\Omega_i$ : pion-isospin-symmetry

$$\Omega_i = \frac{N_{\gamma,i}^{\text{rec,RD}} / N_{\text{ch}}^{\text{rec,RD}}}{N_{\gamma,i}^{\text{rec,MC}} / N_{\text{ch}}^{\text{rec,MC}}}$$

Use  $\Omega_i$  to correct  $\varepsilon_\gamma$ :

$$\varepsilon_\gamma^{\text{MC,corr}}(p_T) = \frac{\sum_i W_i \times dN_{\gamma,i}^{\text{rec,MC}} / dp_T}{dN_\gamma^{\text{prod}} / dp_T}$$



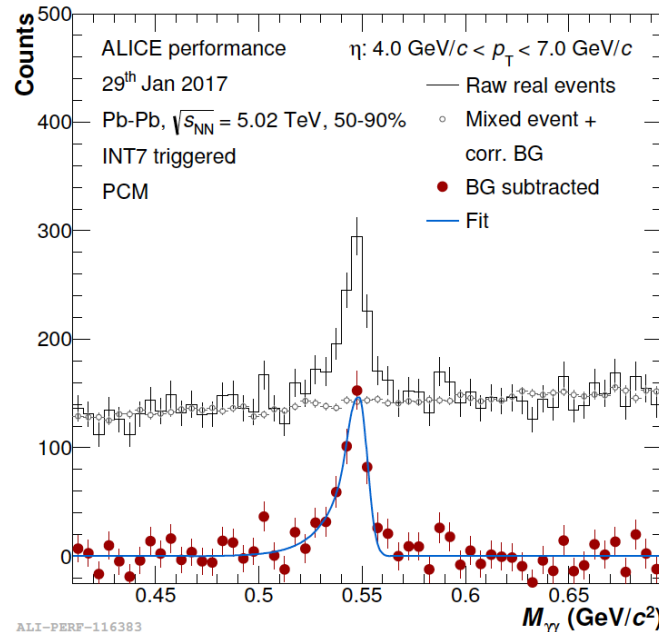
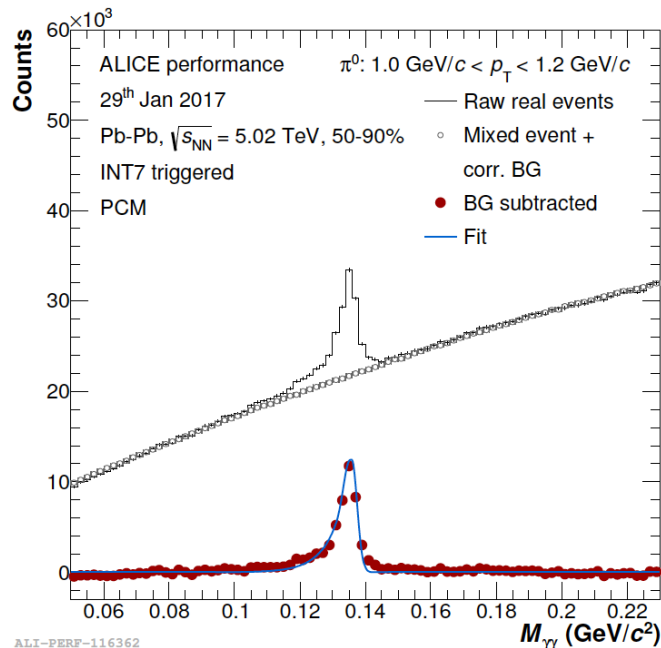
Reduce  $X/X_0$  systematic uncertainty: 4.5%  $\rightarrow$  2.5%  
Mitigate local imperfections in  $X/X_0$  implementation in simulations

# $R_\gamma: \gamma_{\text{inc}}, \text{ neutral mesons and decay photons}$

$$R_\gamma = N_{\gamma,\text{inc}}/N_{\gamma,\text{dec}} \approx \left( \frac{N_{\gamma,\text{inc}}}{\pi^0} \right)_{\text{meas}} / \left( \frac{N_{\gamma,\text{dec}}}{\pi^0} \right)_{\text{sim}}$$

Direct photon signal if  $R_\gamma > 1$

$$N_{\gamma,\text{dir}} = N_{\gamma,\text{inc}} - N_{\gamma,\text{dec}} = \left( 1 - \frac{1}{R_\gamma} \right) \cdot N_{\gamma,\text{inc}}$$



- Measure  $\pi^0$  and  $\eta$  via  $\gamma\gamma$  decay
- Simulation of  $\pi^0, \eta, \omega, \eta'$  decays into  $\gamma$



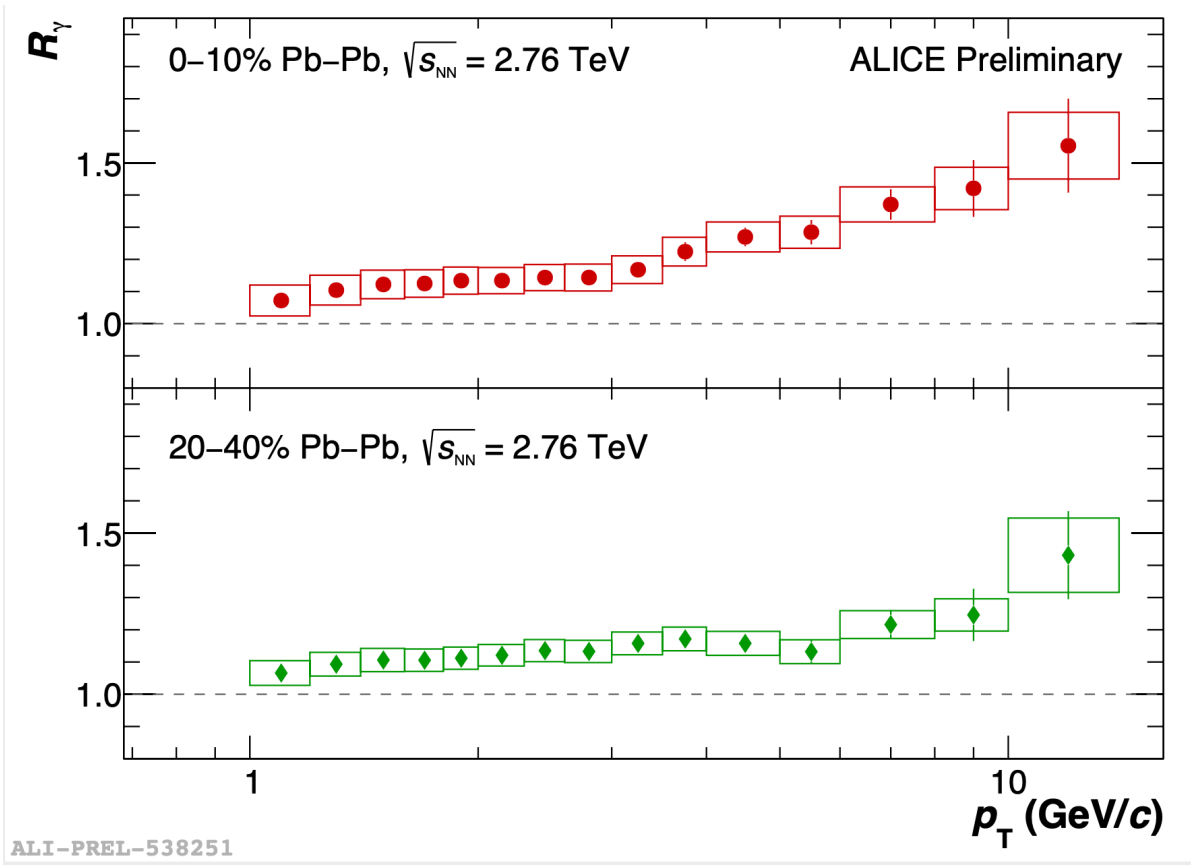
NEW

# Direct photon $R_\gamma$ in Pb – Pb at $\sqrt{s_{NN}} = 2.76$ TeV



Combination of PCM (2011) with  $\Omega_i$  + PHOS (2010)

$$R_\gamma = \frac{\gamma_{inc}}{\pi^0} / \frac{\gamma_{decay}}{\pi^0_{param}} \sim \frac{\gamma_{inc}}{\gamma_{decay}}$$



In agreement with published results  
New centrality available: 0-10%  
Smaller uncertainties

Significant excess for  $p_T > 1$  GeV/c

- 0-10%: 3.1  $\sigma$  ( $1.0 \text{ GeV}/c < p_T < 1.8 \text{ GeV}/c$ )
- 20-40%: 3.4  $\sigma$  ( $1.0 \text{ GeV}/c < p_T < 2.3 \text{ GeV}/c$ )

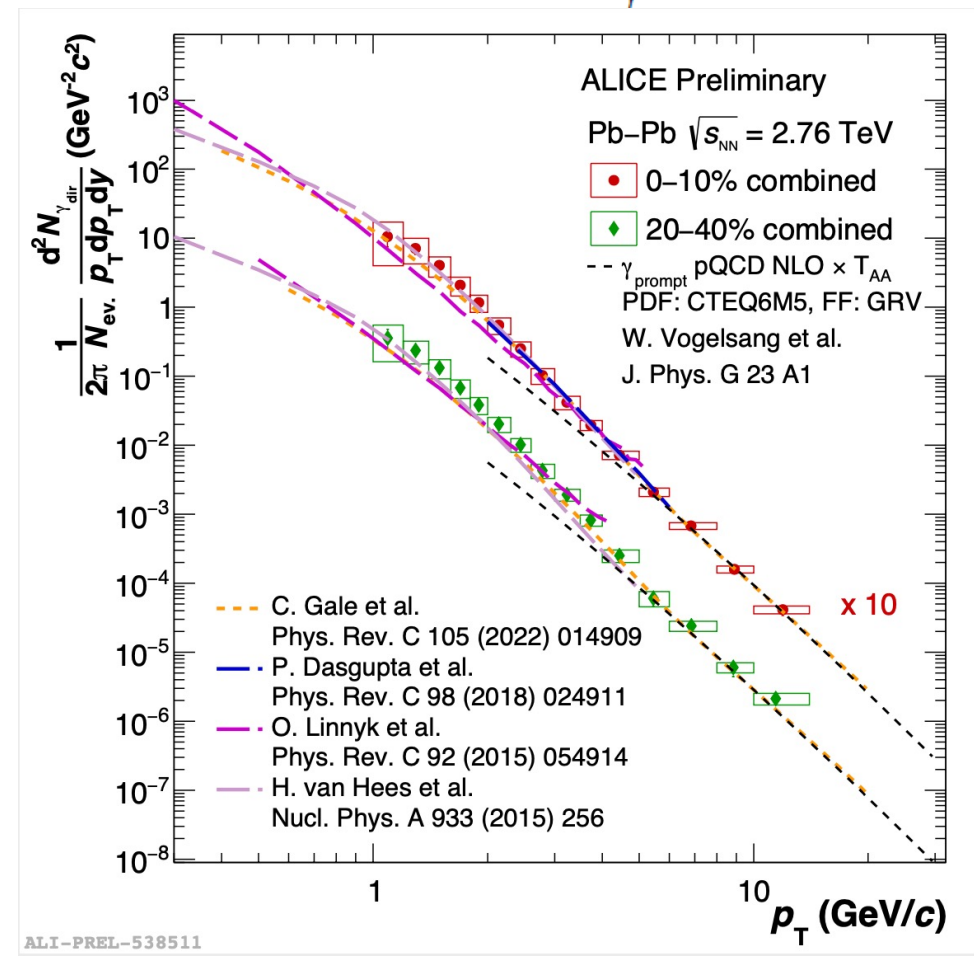
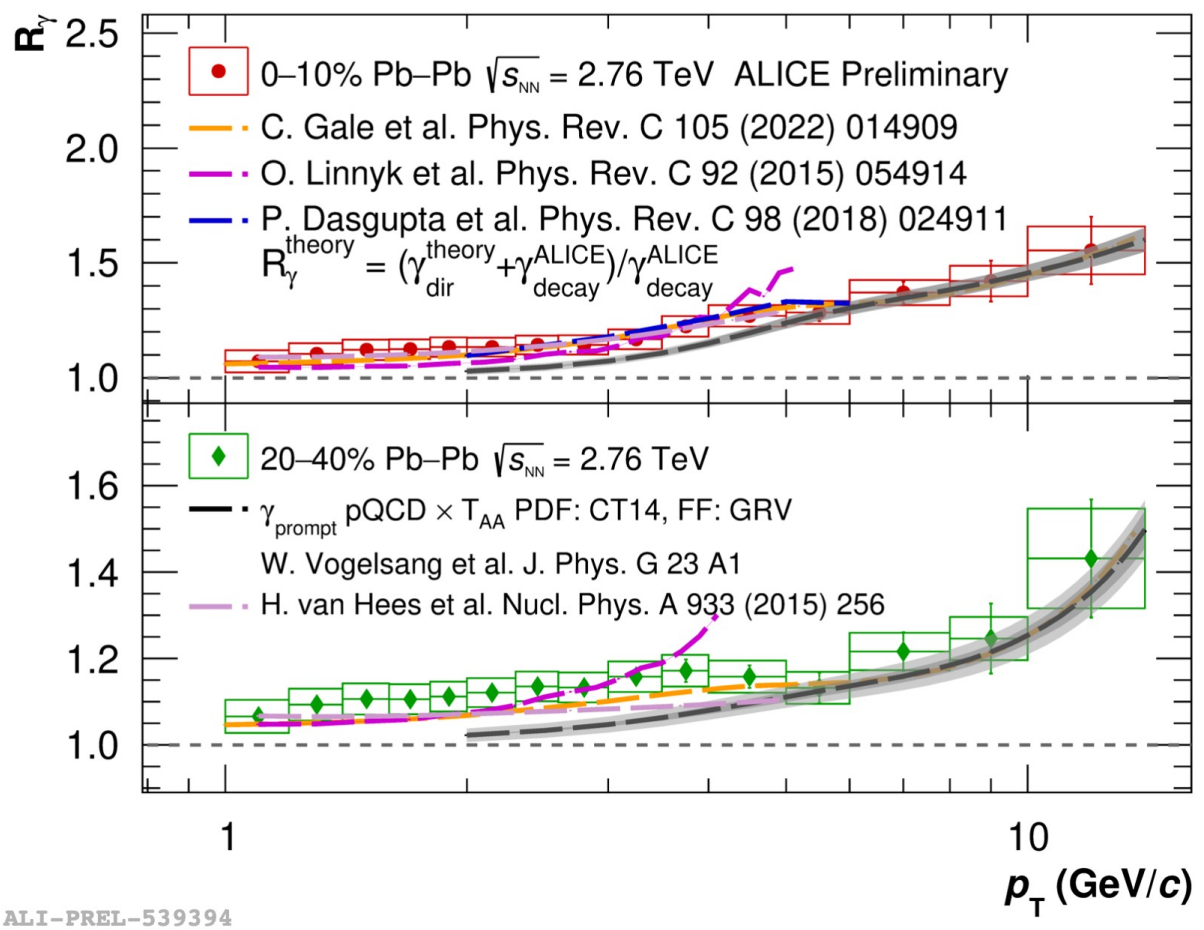
Low  $p_T$  : thermal radiation  
High  $p_T$  : prompt photons

# NEW QGP thermal emission: Pb-Pb at $\sqrt{s_{NN}} = 2.76$ TeV



ALICE

$$N_{\gamma,dir} = \left(1 - \frac{1}{R_{\gamma}}\right) \cdot N_{\gamma,inc}$$

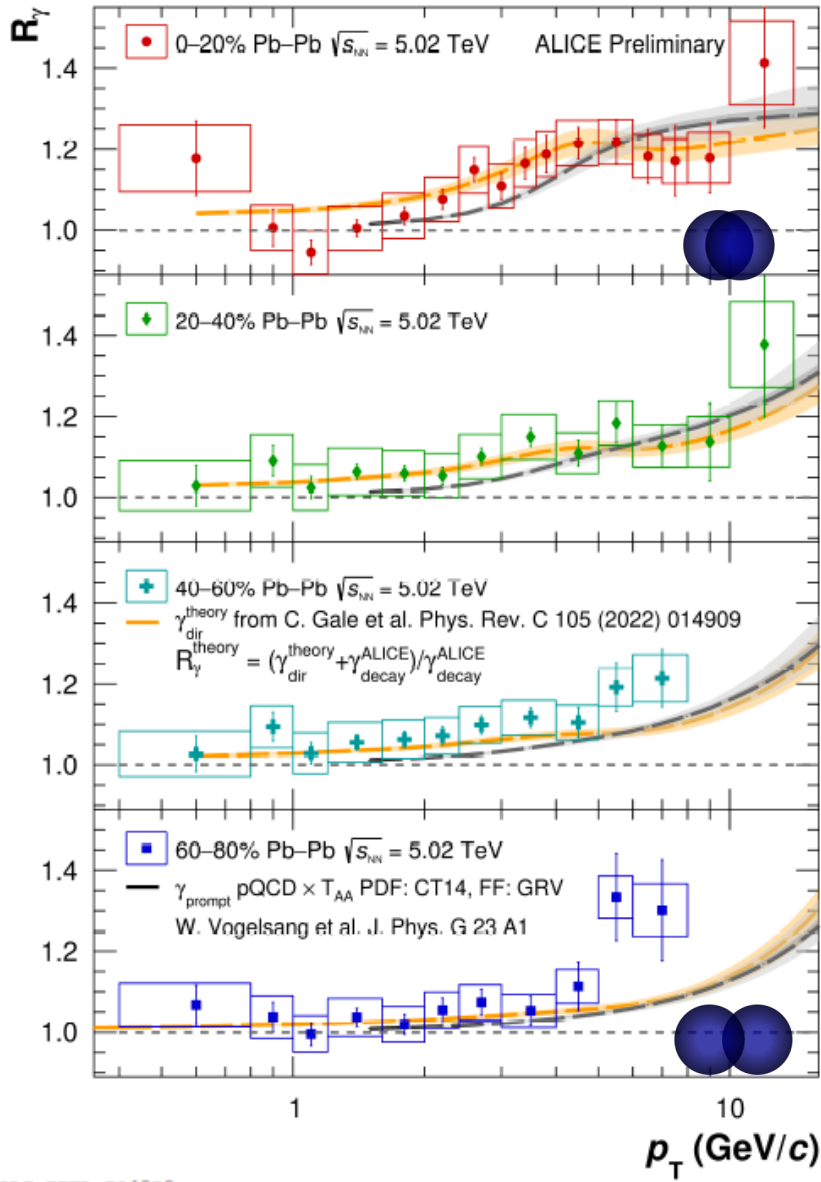


- Excess beyond known prompt yield  $1 < p_T < 4$  GeV/c
- Models that include thermal +(pre-equilibrium) + prompt photons are able to describe the data
- Not yet possible to discriminate among different models

# QGP thermal emission: Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV



ALICE



$$R_\gamma = \frac{\gamma_{inc}}{\pi^0} / \frac{\gamma_{decay}}{\pi^0_{param}} \sim \frac{\gamma_{inc}}{\gamma_{decay}}$$

- At low  $p_T$ :  
 $R_\gamma$  is close to 1  $\rightarrow$  small thermal contribution
- For  $p_T > 2-3$  GeV/c:  
 Excess  $\rightarrow$  pre-equilibrium and prompt photons
- Data consistent with NLO pQCD calculation of prompt photons  $\times T_{AA}$   
 Calculation by W. Vogelsang, using PDF: CT14, FF: GRV
- Thermal+ pre-equilibrium photons + prompt photon:  $R_\gamma \sim 1.05 \rightarrow$   
 Better data description better than with only prompt photons  
 IP-Glasma initial conditions + K $\emptyset$ MP $\emptyset$ ST+ MUSIC viscous hydrodynamics,  
 prompt  $\gamma$  PDF:nCTEQ15-np, FF: BFG-II

Bands represent (theoretical and) experimental uncertainties



# QGP thermal emission: Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV

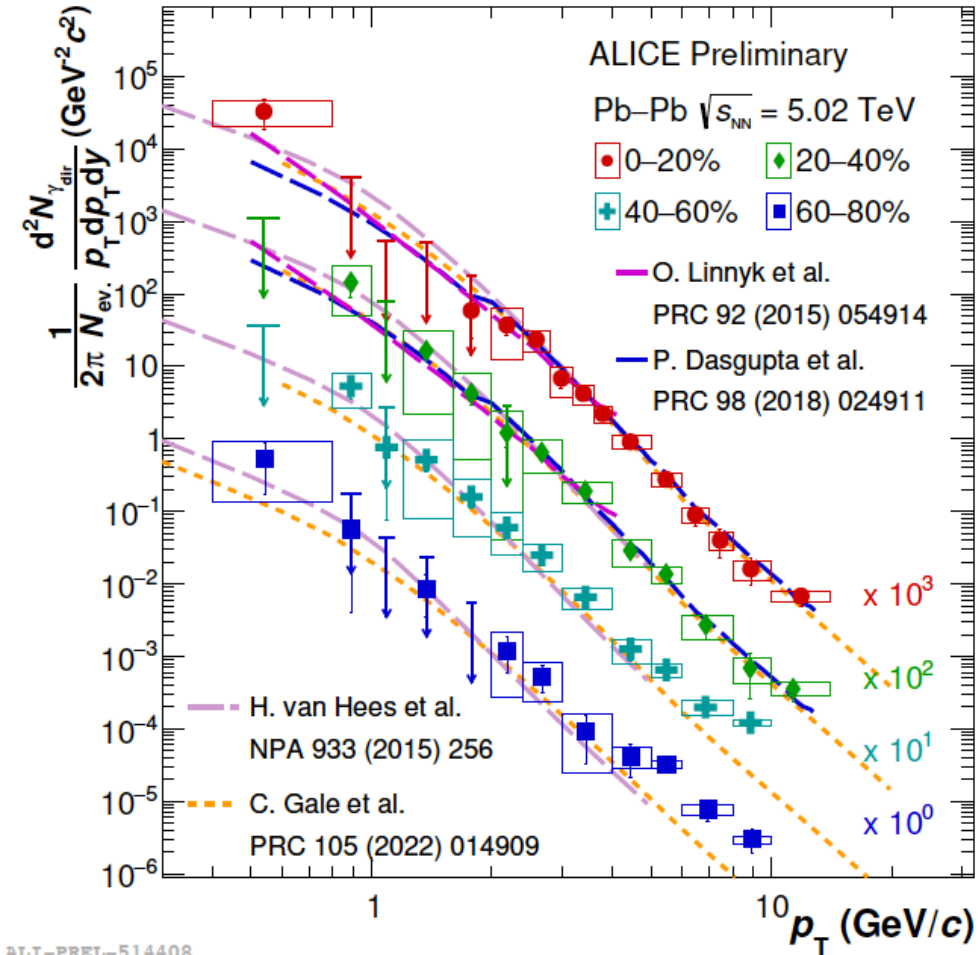


ALICE

$$N_{\gamma,dir} = \left(1 - \frac{1}{R_\gamma}\right) \cdot N_{\gamma,inc}$$

- Upper limits (90% CL) given where  $N_{\gamma,dir}$  consistent with 0
  - Different model calculations of direct photons:
    - Microscopic transport approach (PHSD)
    - Relativistic hydrodynamic, different initial conditions, thermalization times, hadronization temperatures, with and without pre-equilibrium  $\gamma$
  - At high  $p_T$  consistent with pQCD
- Not yet possible to favor a model over the other

Expect more precise results with the full Run 2 data and Run 3

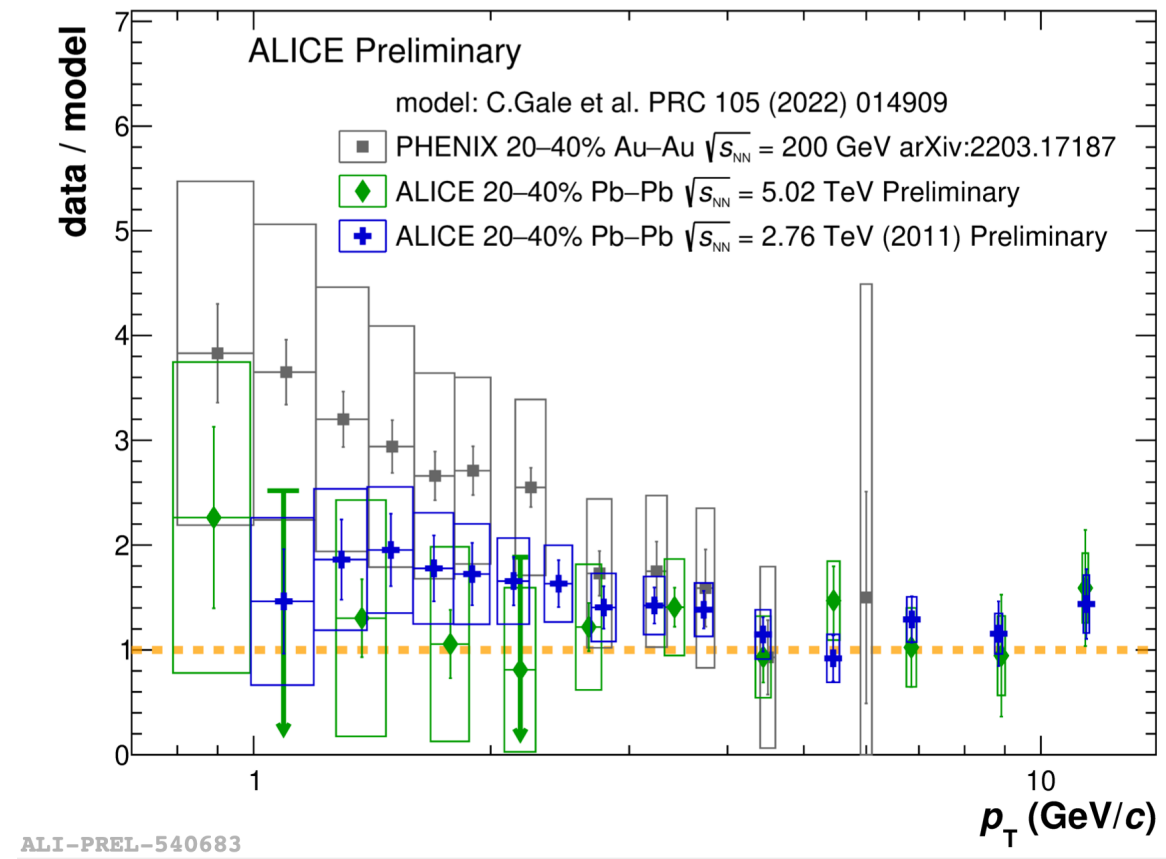
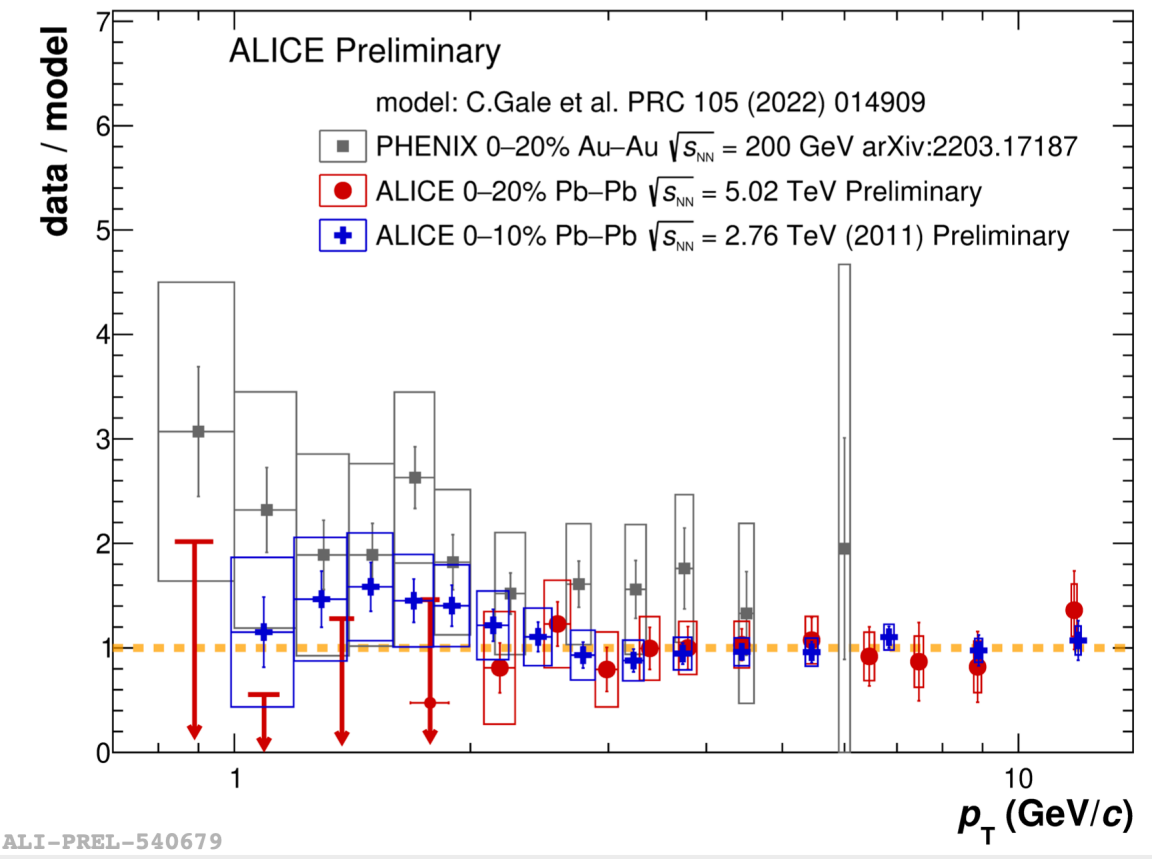


ALI-PREL-514408

R. Bailhache, talk 44

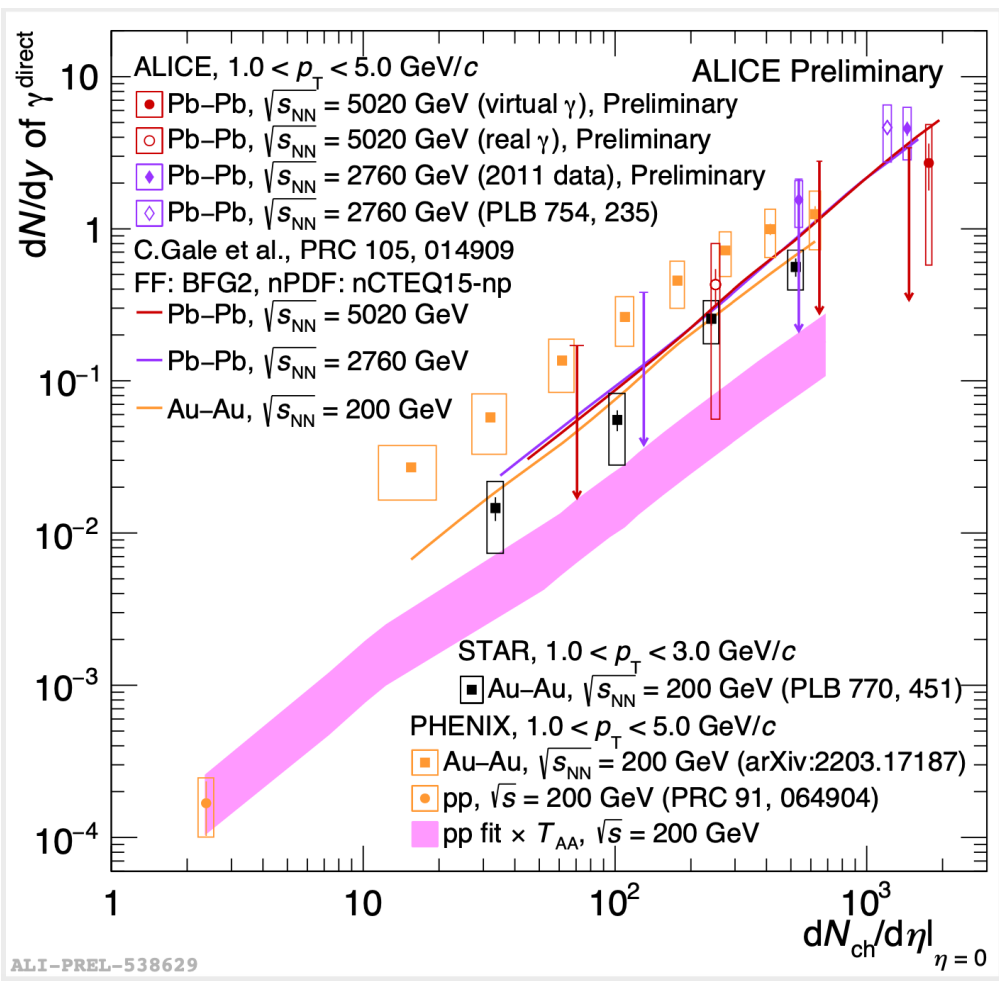
# Direct photon puzzle in yields?

Ratio between direct photon production and their respective state-of-the-art model calculation



Good agreement between ALICE data and model predictions  
 Slight tension at low  $p_T$  for the PHENIX data  
 Future: puzzle involving direct photon flow?

# Integrated direct photon yield vs $dN_{ch}/d\eta$



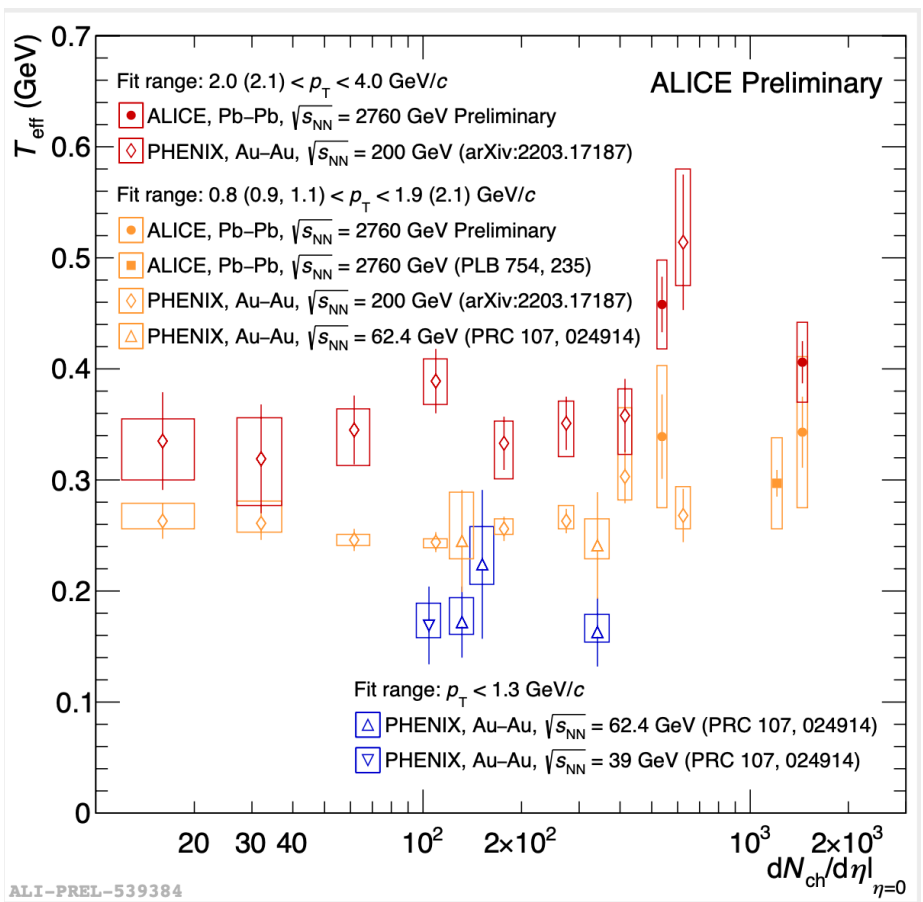
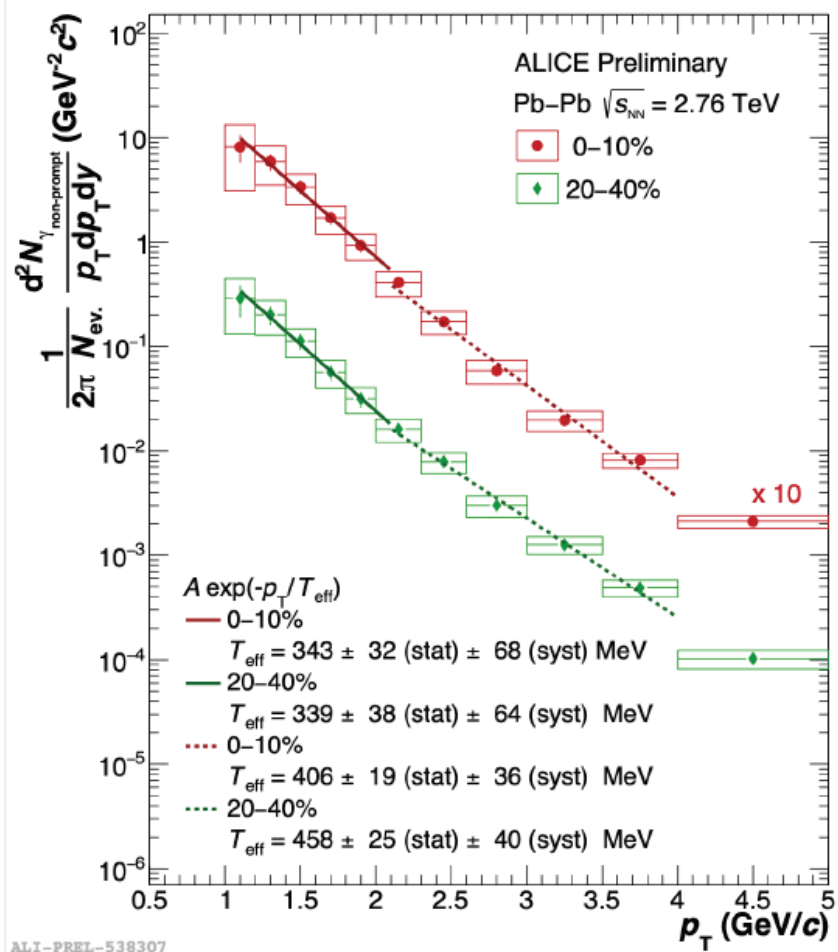
- Integrated direct photon yield ( $1 < p_T < 5$  GeV/c) vs  $dN_{ch}/d\eta$
- ALICE data points follow similar trend as observed in PHENIX and STAR experiments and as predicted by hydro model

Universal power-law scaling of direct  $\gamma$  yield vs  $N_{ch}$  seen for different systems and collision energies

# T<sub>eff</sub> from non-prompt photons

Non-prompt  $\gamma$  = direct  $\gamma$  - T<sub>AA</sub> · pQCD

is T<sub>eff</sub> (2.1 < p<sub>T</sub> < 4 GeV/c) > T<sub>eff</sub> (1.1 < p<sub>T</sub> < 2.1 GeV/c) ?  
pre-equilibrium photons? earlier time emission?



# Bose-Einstein $\gamma\gamma$ correlations in Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV

HBT interferometry (two-particle correlations)

Space – time dimensions of emitting source

$\gamma\gamma$ -HBT  $\rightarrow$  source size  
direct photon fraction

Correlation function:

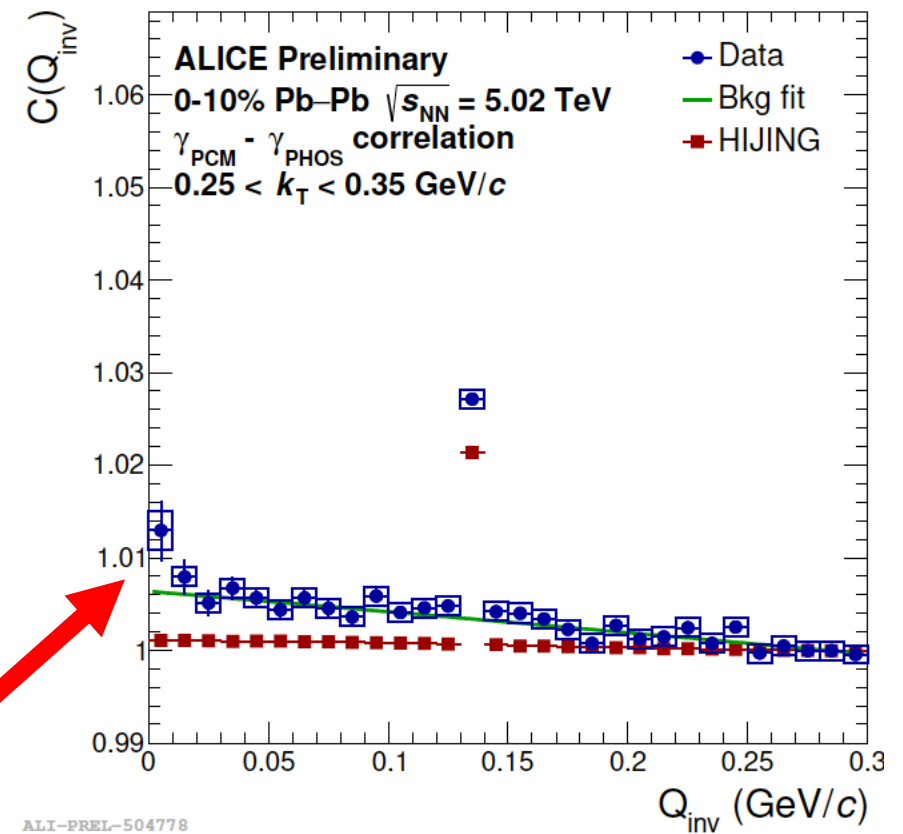
$$\gamma_{\text{PHOS}} - \gamma_{\text{PCM}}$$

$$C(Q_{\text{inv}}) = A(Q_{\text{inv}})/B(Q_{\text{inv}})$$

Bins of  $k_{\text{T}}$  (average pair momentum) and centrality

Small hint of a HBT-like effect at lower  $Q_{\text{inv}}$

$$C(Q_{\text{inv}}) = 1 + \lambda_{\text{inv}} \exp(-R_{\text{inv}}^2 Q_{\text{inv}}^2)$$

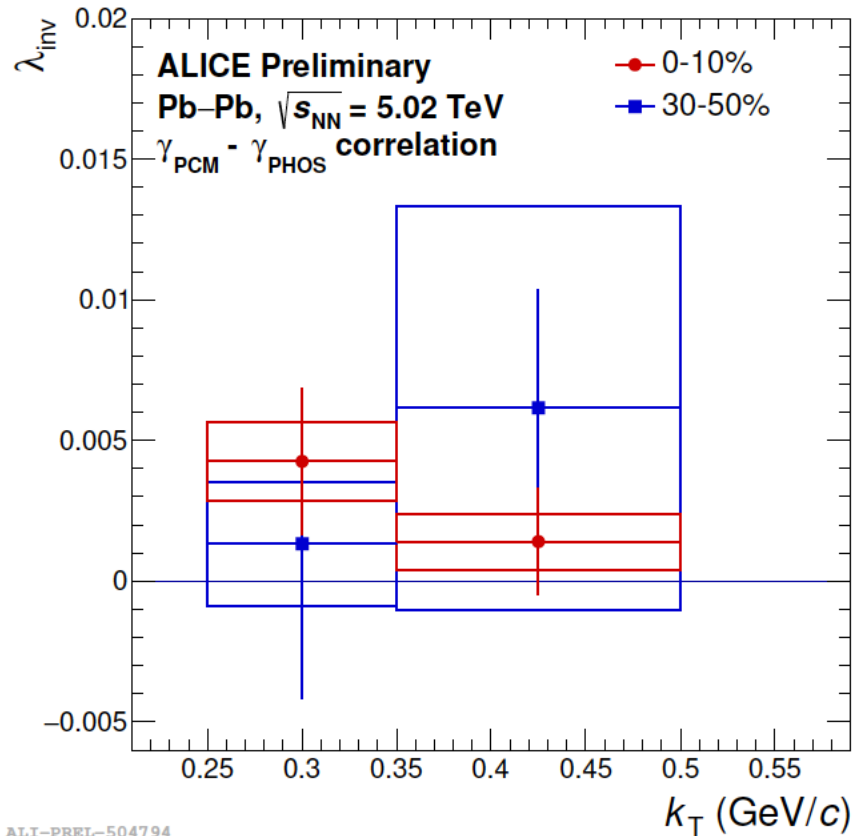




# Bose-Einstein $\gamma\gamma$ correlations in Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV



Sensitive to the source size and to the direct photon fraction



$$C(Q_{inv}) = 1 + \lambda_{inv} \exp(-R_{inv}^2 Q_{inv}^2)$$

$\lambda_{inv}$  not significantly different from zero

$$r_\gamma = \frac{N_{dir}}{N_{inc}} = \sqrt{2\lambda}$$

[WA98, PRL 93,022301 \(2004\)](#)

Ongoing: Measurements performed in the LCMS

# Summary

- Direct photon production in Pb-Pb collisions with improved  $X/X_0$  uncertainties
  - at  $\sqrt{s_{NN}} = 2.76$  TeV
    - Significant excess for  $p_T > 1$  GeV/c
    - $T_{\text{eff}}$  as function of  $dN_{\text{ch}}/d\eta$  extracted in two  $p_T$  ranges. Consistent values at similar  $dN_{\text{ch}}/d\eta$
  - at  $\sqrt{s_{NN}} = 5.02$  TeV
    - Significant excess of prompt photons at  $p_T > 3$  GeV/c
    - $R_\gamma$  at lower  $p_T$  consistent with unity

Integrated direct  $\gamma$  yields follow power law scaling with  $dN_{\text{ch}}/d\eta$

Model calculations consistent with the data, no yet possible to discriminate

- Photon HBT provides a complementary method to obtain  $R_\gamma$ , and possibly the source size
- Stay tuned for results with full Run 2 statistics and Run 3 data

# Thank you

# Backup slides

# Data-driven precision determination of the material budget in ALICE

[arxiv: 2303.15317](https://arxiv.org/abs/2303.15317)

$$\omega_i = \frac{P_i^{\text{RD}} \times \epsilon_{\gamma,i}^{\text{RD}} \times \epsilon_{\gamma,\text{gas}}^{\text{MC}}}{P_i^{\text{MC}} \times \epsilon_{\gamma,i}^{\text{MC}} \times \epsilon_{\gamma,\text{gas}}^{\text{RD}}},$$

$$\Omega_i = \frac{P_i^{\text{RD}} \times \epsilon_{\gamma,i}^{\text{RD}} \times \epsilon_{\text{track}}^{\text{MC}}}{P_i^{\text{MC}} \times \epsilon_{\gamma,i}^{\text{MC}} \times \epsilon_{\text{track}}^{\text{RD}}}.$$

$$\frac{\Omega_i}{\omega_i} = \frac{\epsilon_{\gamma,\text{gas}}^{\text{RD}} \times \epsilon_{\text{track}}^{\text{MC}}}{\epsilon_{\gamma,\text{gas}}^{\text{MC}} \times \epsilon_{\text{track}}^{\text{RD}}}$$

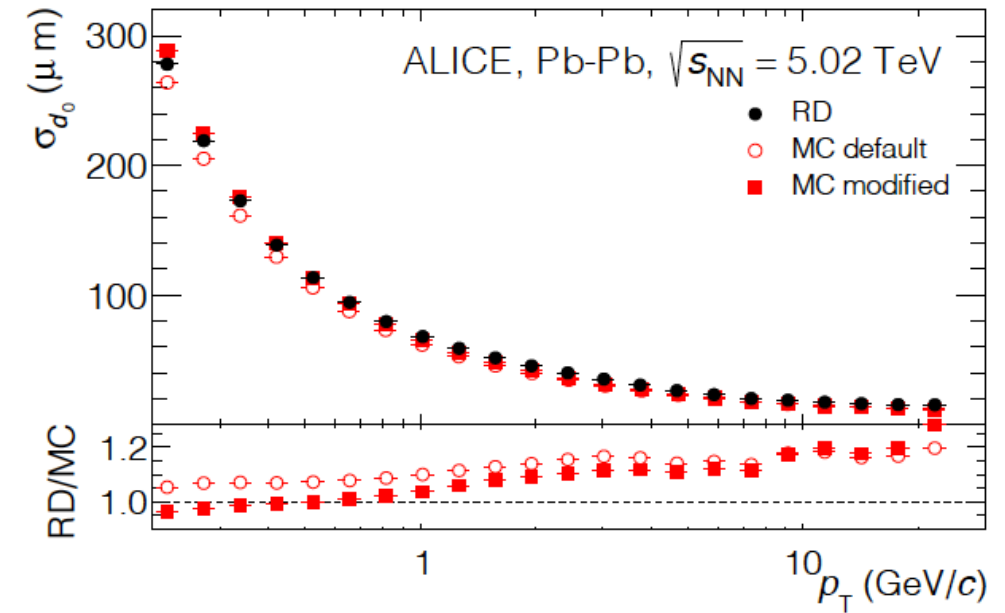
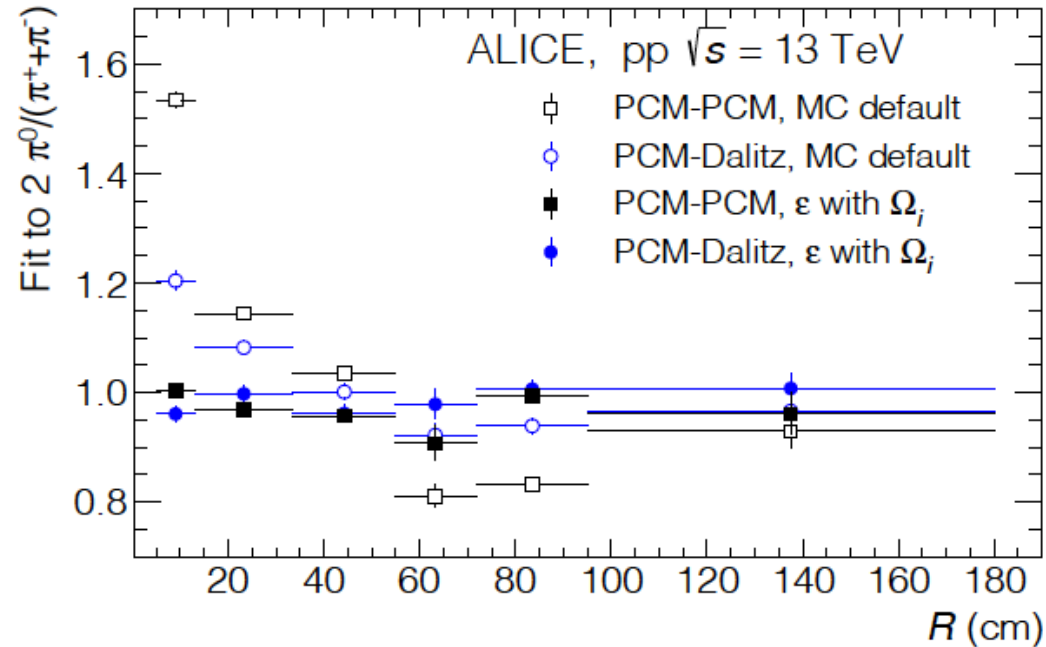
	$\Omega_i$		$\omega_i$	
	5 cm < R < 8.5 cm	95 cm < R < 145 cm	8.5 cm < R < 13 cm	72 cm < R < 95 cm
<b>V<sup>0</sup> finder</b>	2.74 %	2.9%	2.2%	1.83%
<b>Generator</b>	0.16%	2.9%	3.2 %	0.62 %
$p_{T,\text{min}}$	Negligible	Negligible	Negligible	Negligible
$\sigma_{\text{sys}}$	2.74%	4.1%	3.8%	1.93%

R interval	R range (cm)	$\Omega_i$	$\sigma_{\text{stat}}$ %	$\sigma_{\text{sys}}$ %	$\sigma_{\text{total}}$ %
0	0–1.5	0.9859	1.2	-	-
1	1.5–5	1.177	0.42	-	-
2	5–8.5	1.240	0.36	2.7	2.8
3	8.5–13	1.238	0.42	0.77	0.9
4	13–21	1.067	0.34	2.0	2.1
5	21–33.5	1.081	0.25	1.7	1.7
6	33.5–41	1.039	0.35	3.1	3.1
7	41–55	1.001	0.30	1.5	1.5
8	55–72	0.926	0.35	3.7	3.7
9	72–95	0.943	0.19	3.7	3.7
10	95–145	0.975	0.62	4.1	4.1
11	145–180	0.932	0.89	1.4	1.6
<b>average</b>	<b>5–180</b>	<b>1.04</b>	<b>0.312%</b>	<b>2.5%</b>	<b>2.5%</b>

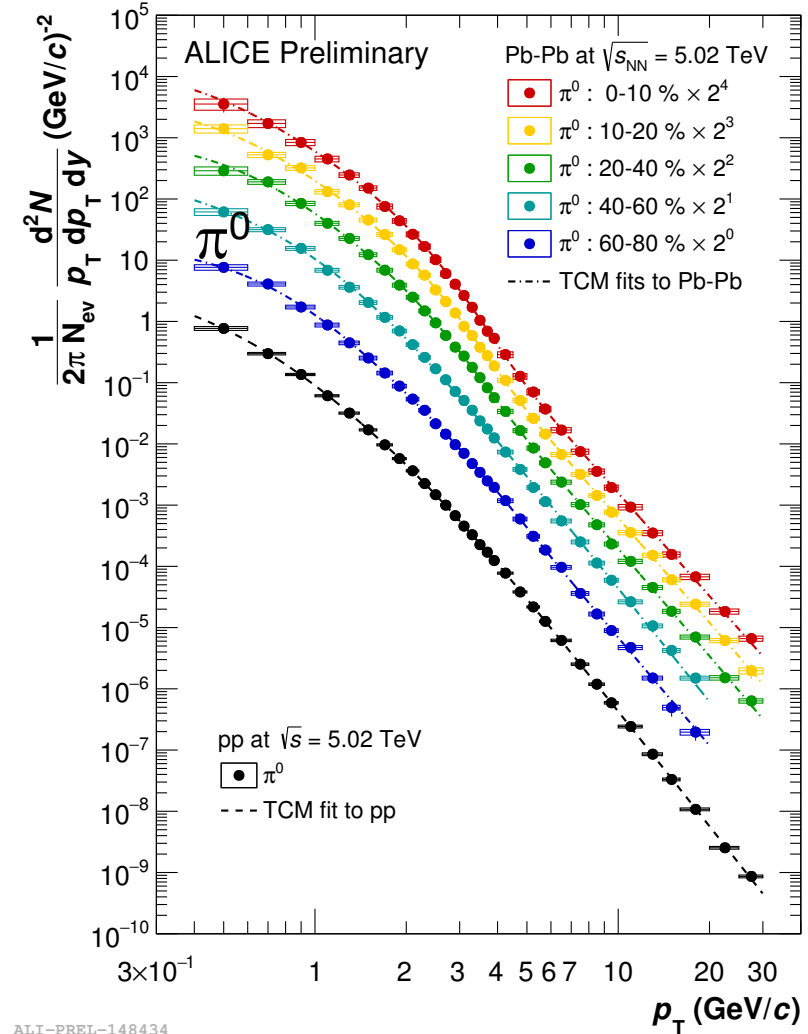
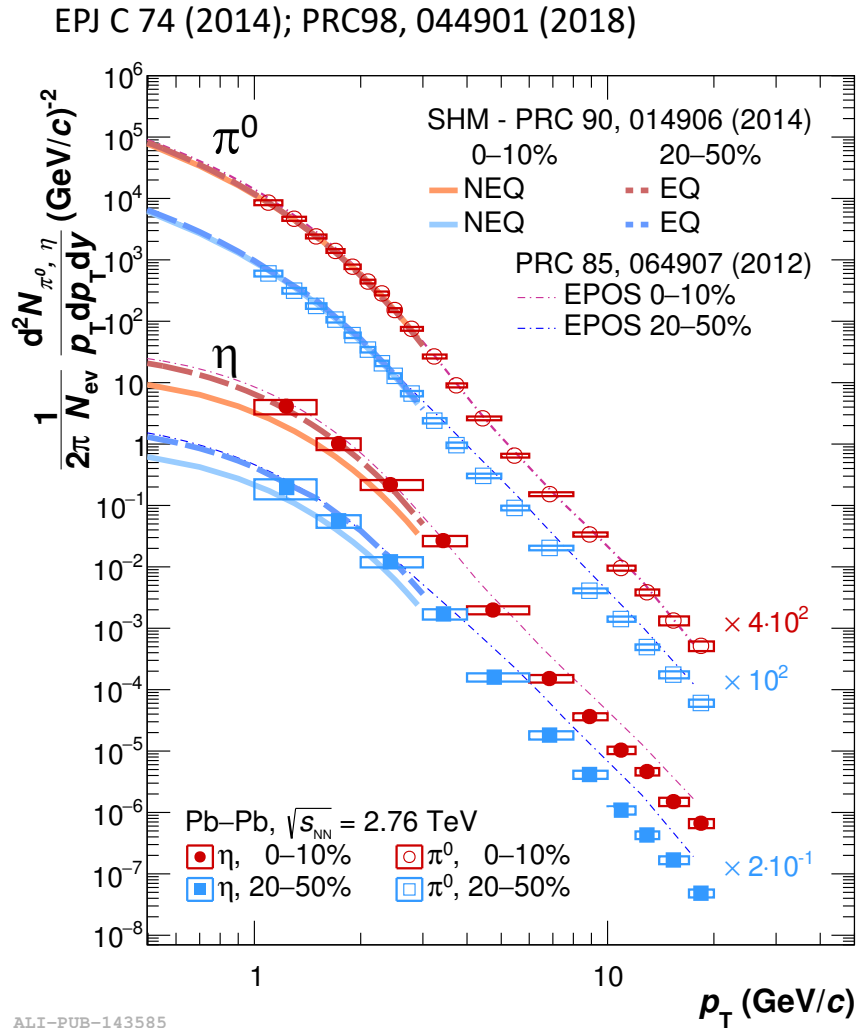


# Data-driven precision determination of the material budget in ALICE

[arxiv: 2303.15317](https://arxiv.org/abs/2303.15317)



# $\pi^0, \eta$ : Pb-Pb at $\sqrt{s_{NN}} = 2.76, 5.02$ TeV



First  $\eta$  measurement in Pb-Pb at the LHC

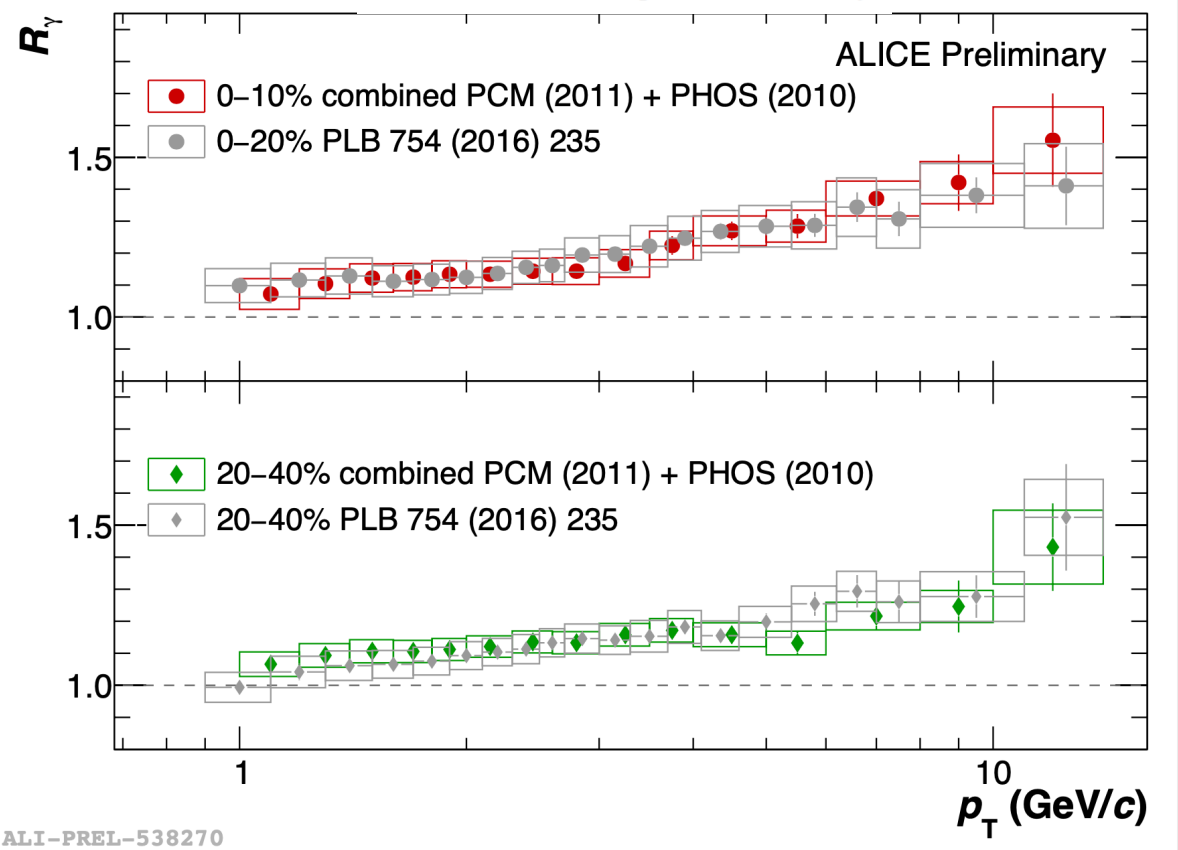
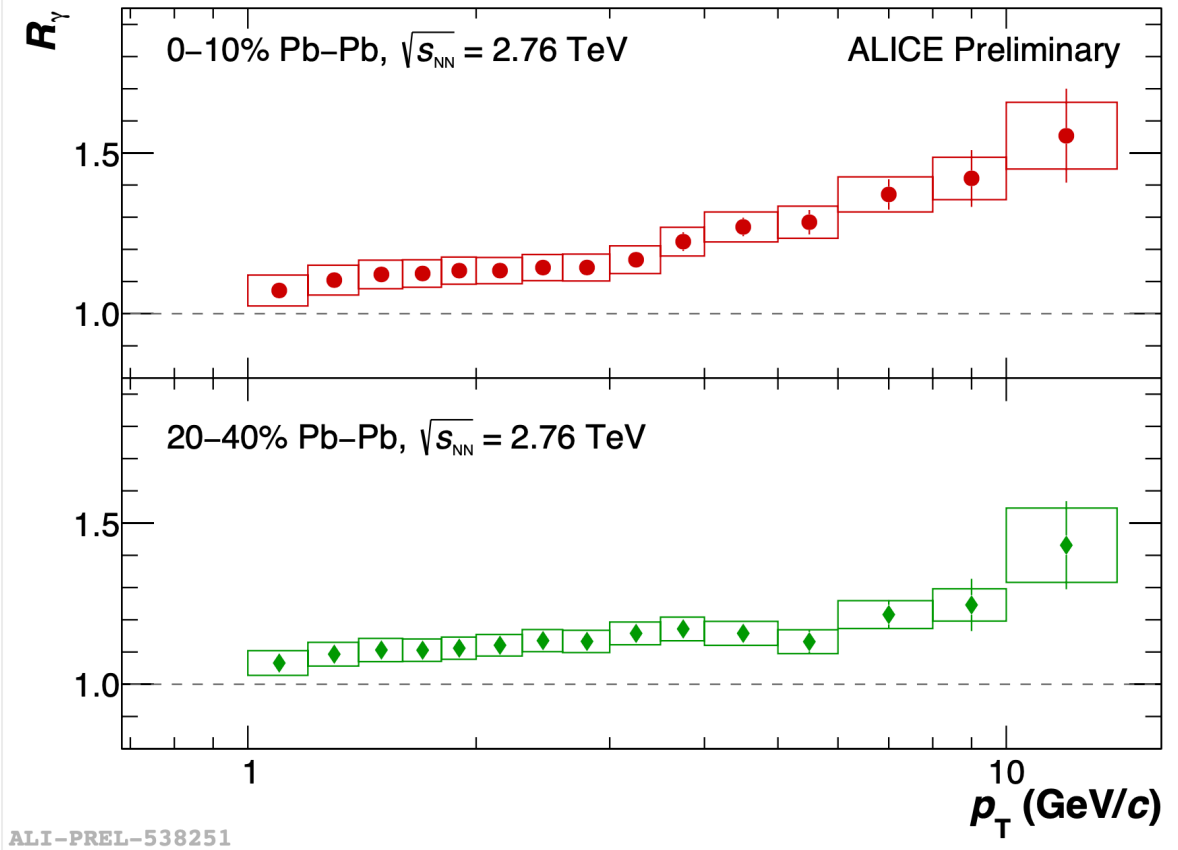
NEW

# Direct photon $R_\gamma$ in Pb – Pb $\sqrt{s_{NN}} = 2.76$ TeV



Combination of PCM (2011) with  $\Omega_i$  + PHOS (2010)

$$R_\gamma = \frac{\gamma_{inc}}{\pi^0} / \frac{\gamma_{decay}}{\pi^0_{param}} \sim \frac{\gamma_{inc}}{\gamma_{decay}}$$



Significant excess for  $p_T > 1$  GeV/c

Low  $p_T$ : thermal radiation  
 High  $p_T$ : prompt photons

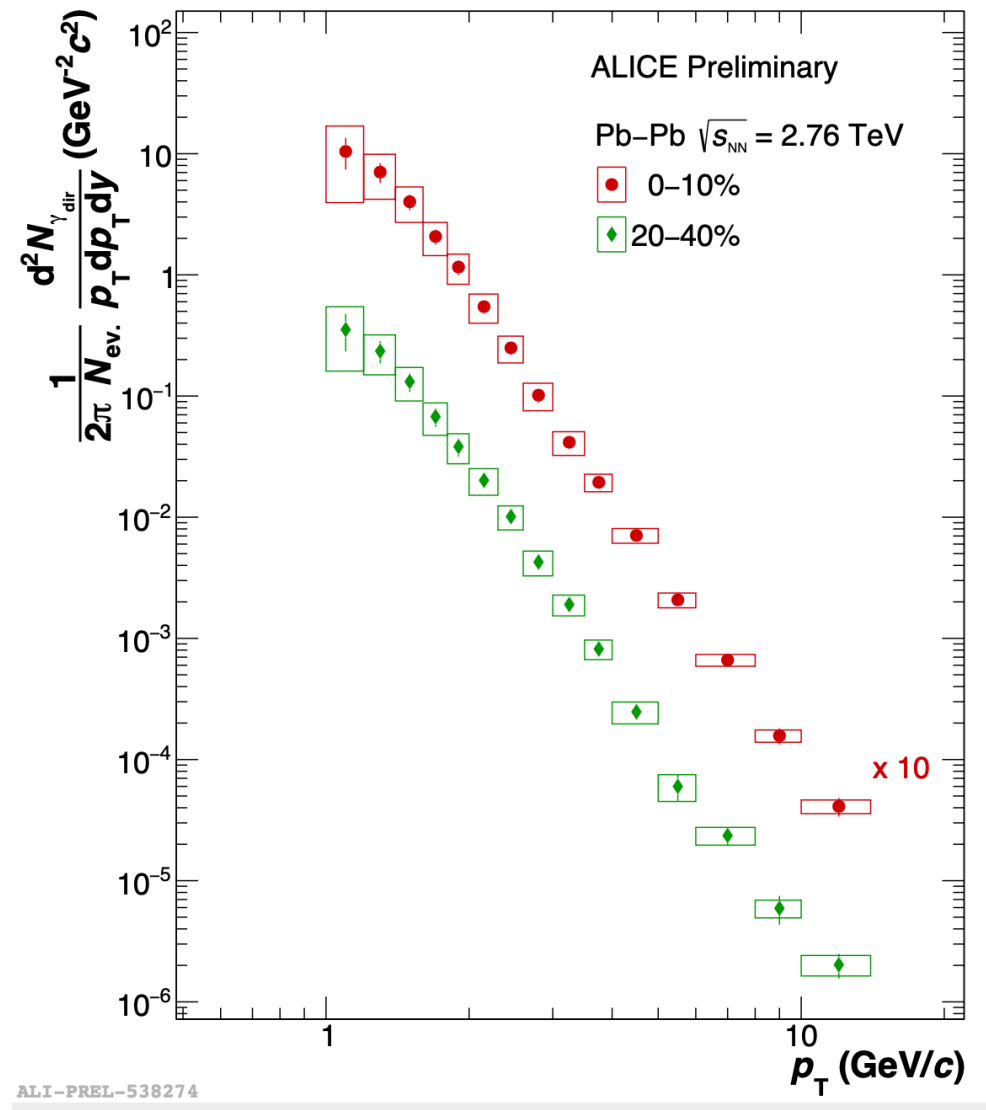
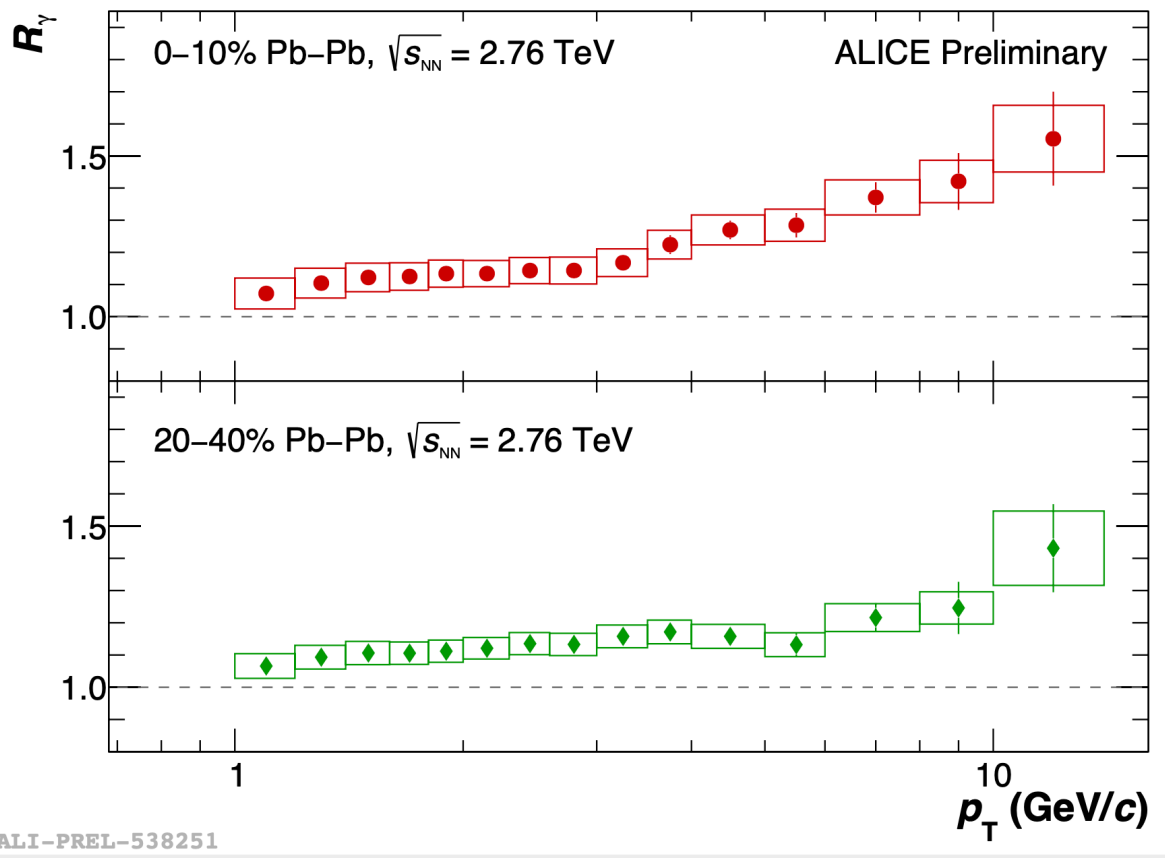
In agreement with published results  
 New centrality available: 0-10%  
 Smaller uncertainties

NEW

# Direct photon $R_\gamma$ in Pb – Pb $\sqrt{s_{NN}} = 2.76$ TeV



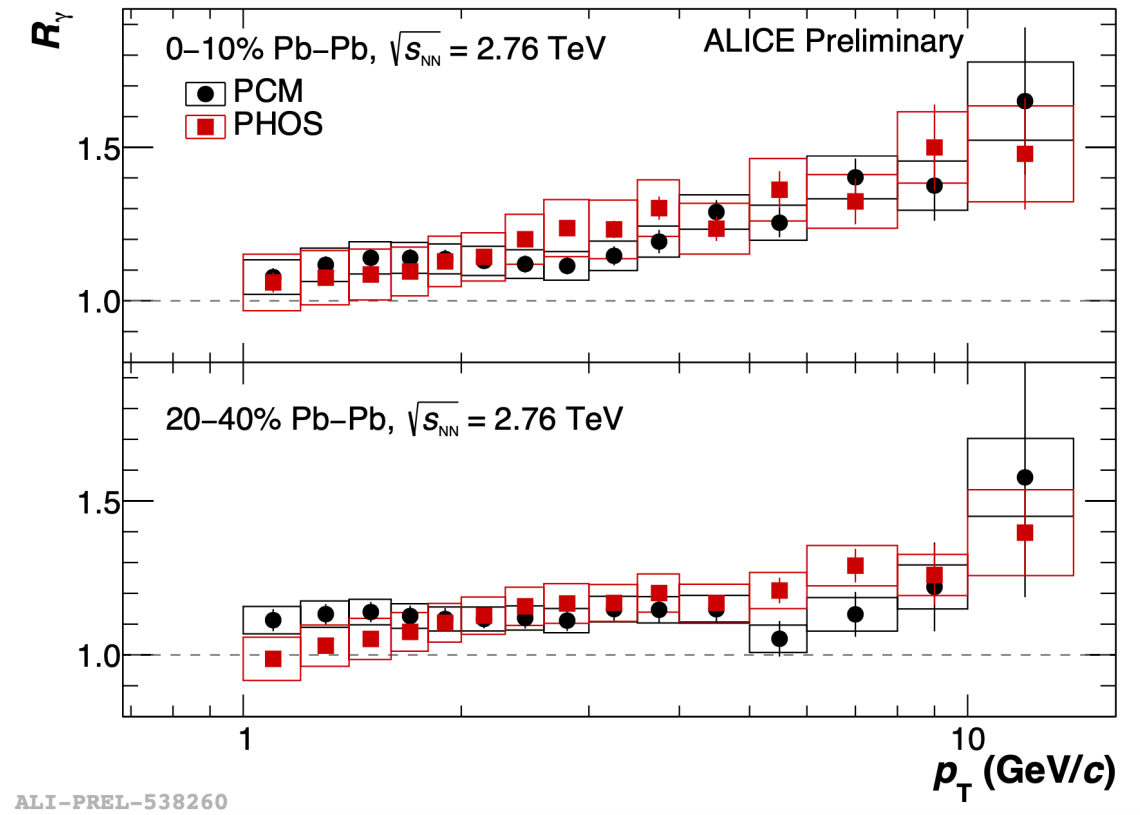
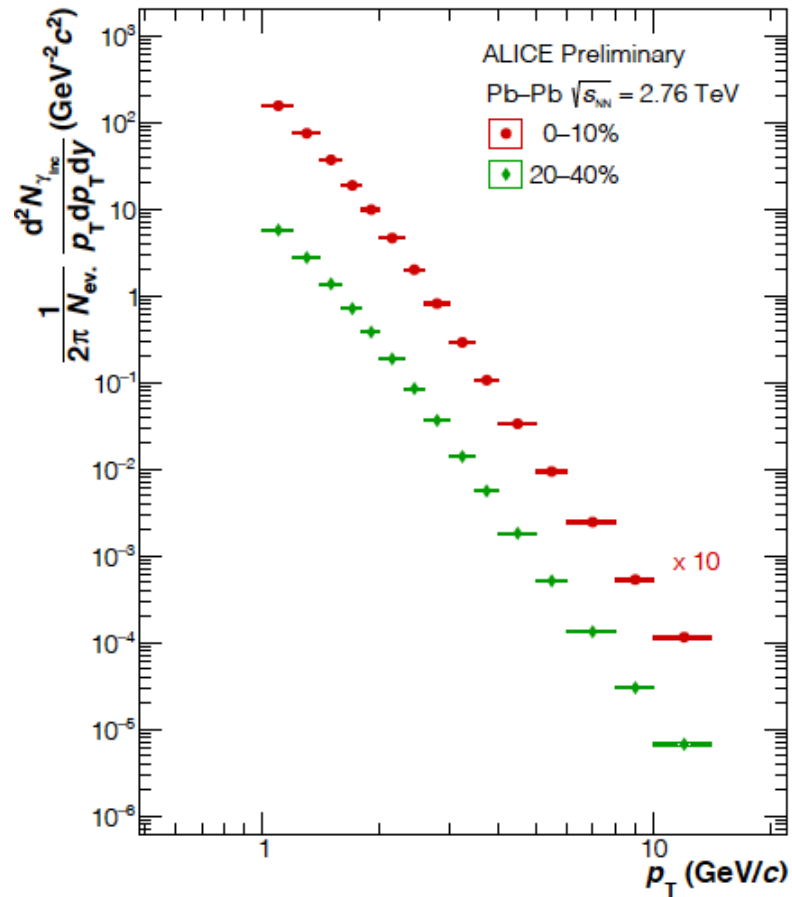
Combination of PCM(2011) with  $\Omega_i$ +PHOS(2010)



Significant excess for  $p_T > 1$  GeV/c

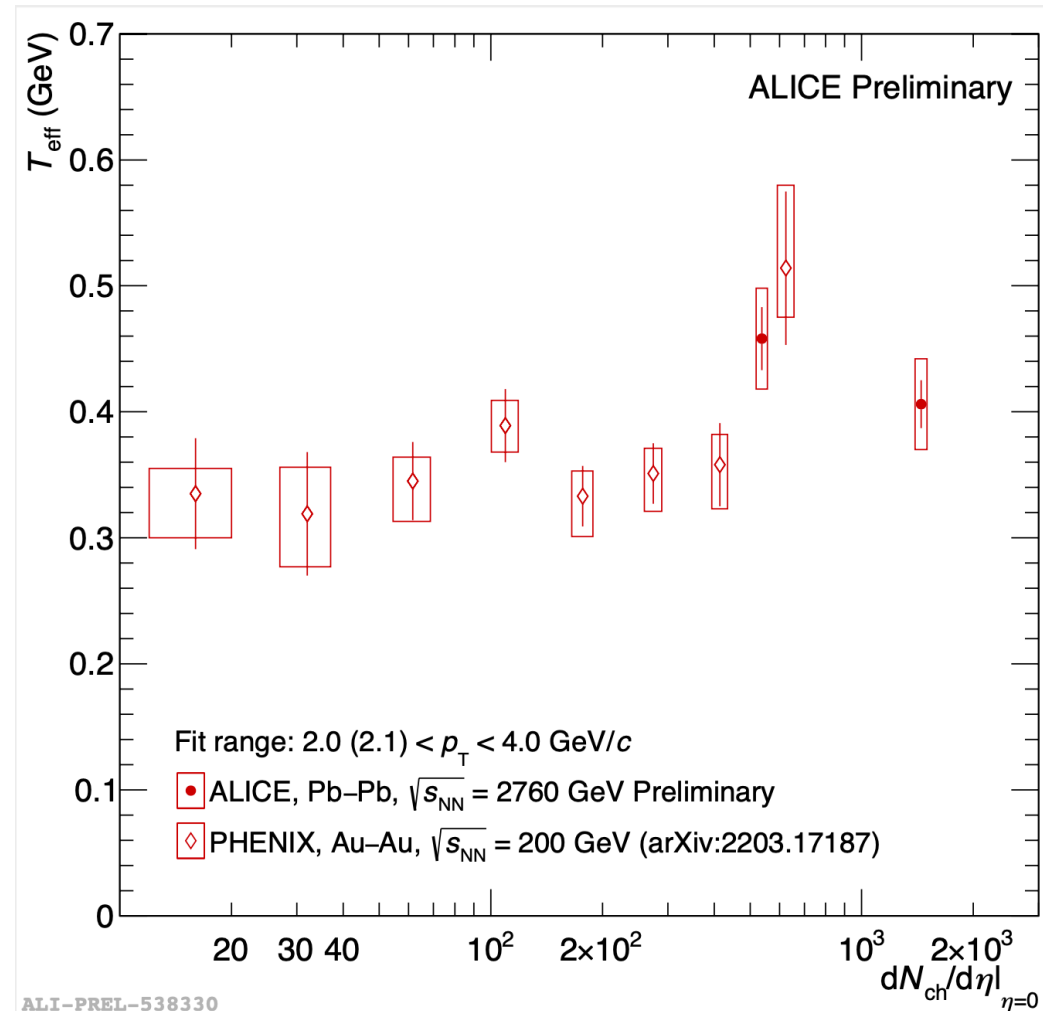
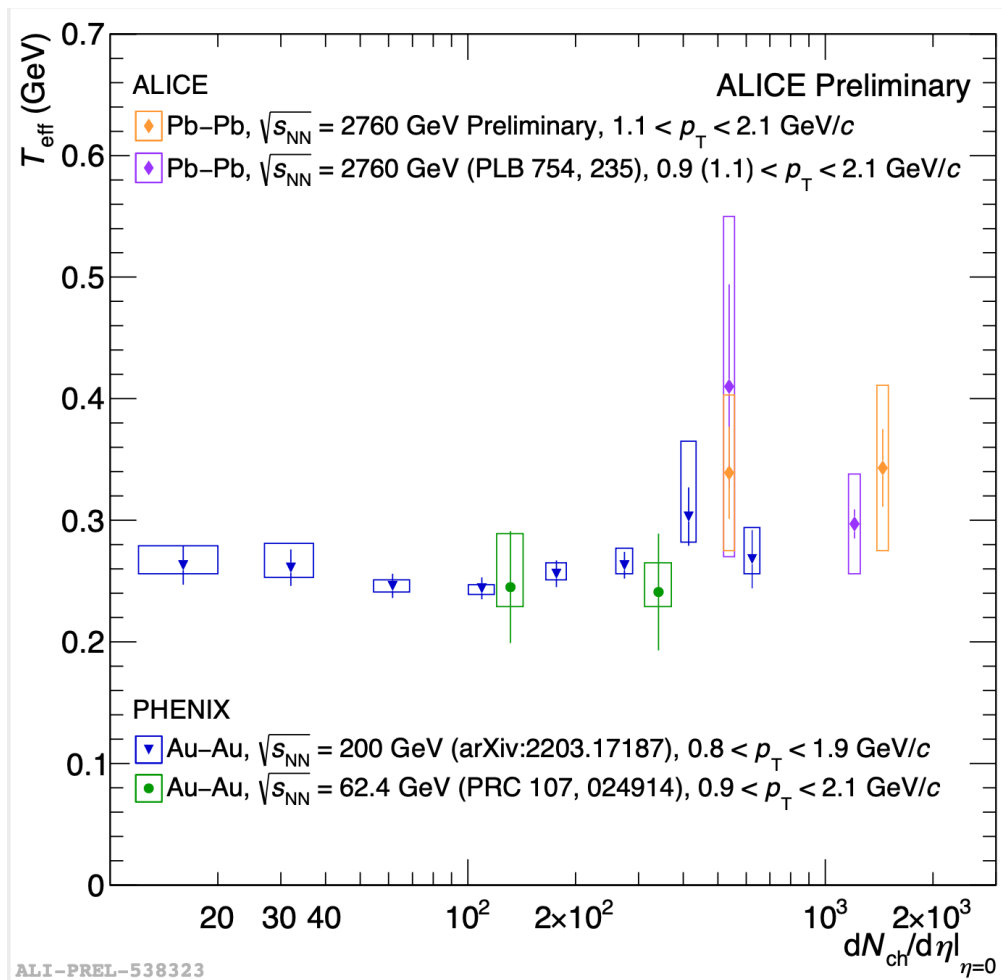
Low  $p_T$  : thermal radiation  
High  $p_T$  : prompt photons

# Inclusive $\gamma$ and $R_\gamma$ in Pb-Pb 2.76 TeV

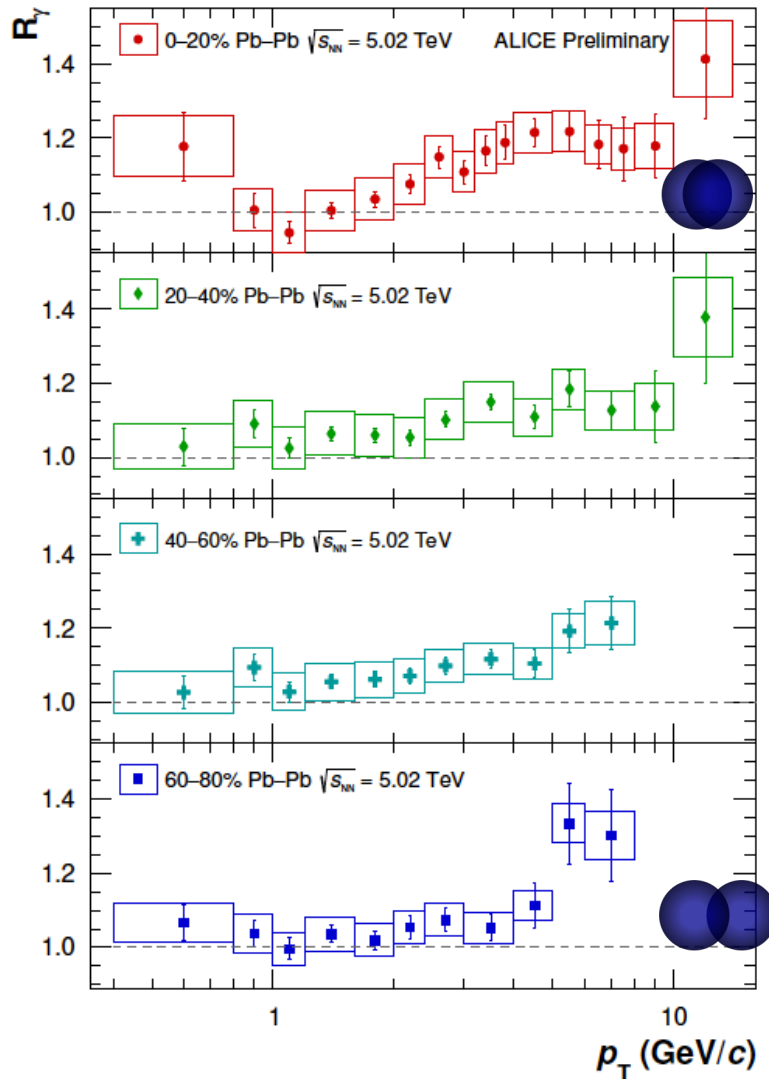




# $T_{\text{eff}}$ from non-prompt photons



# Direct photon $R_\gamma$ in Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV



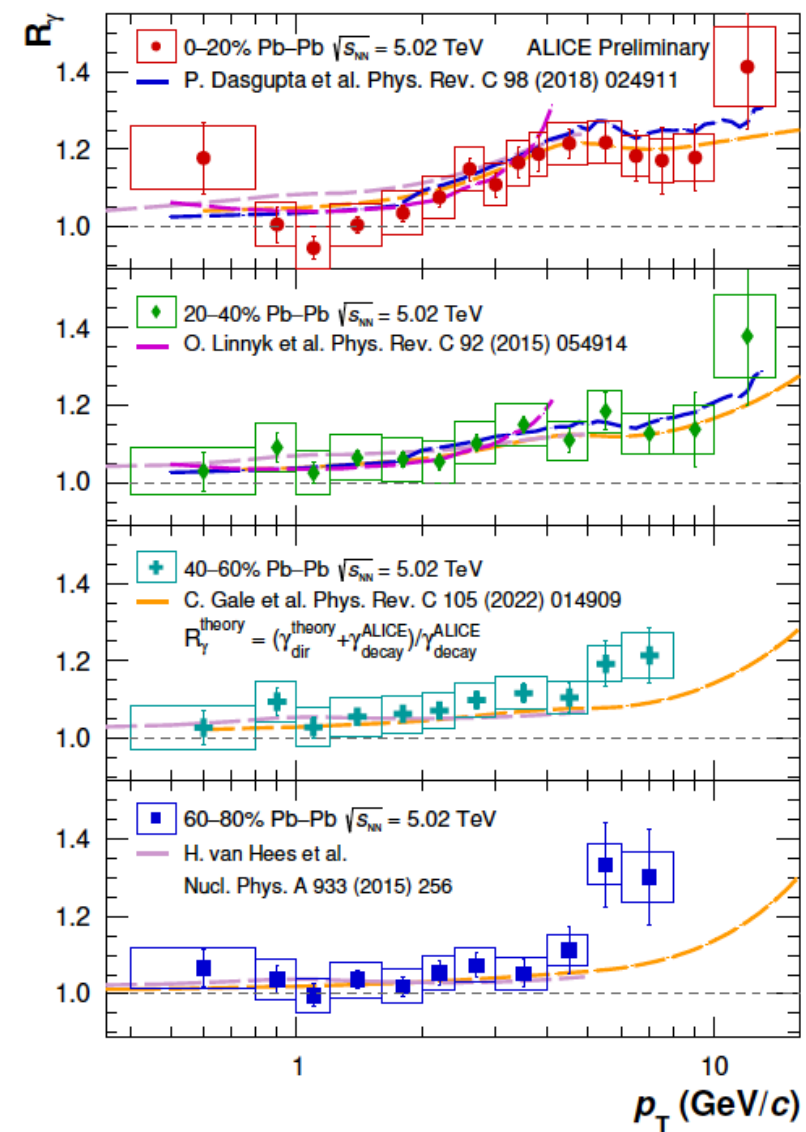
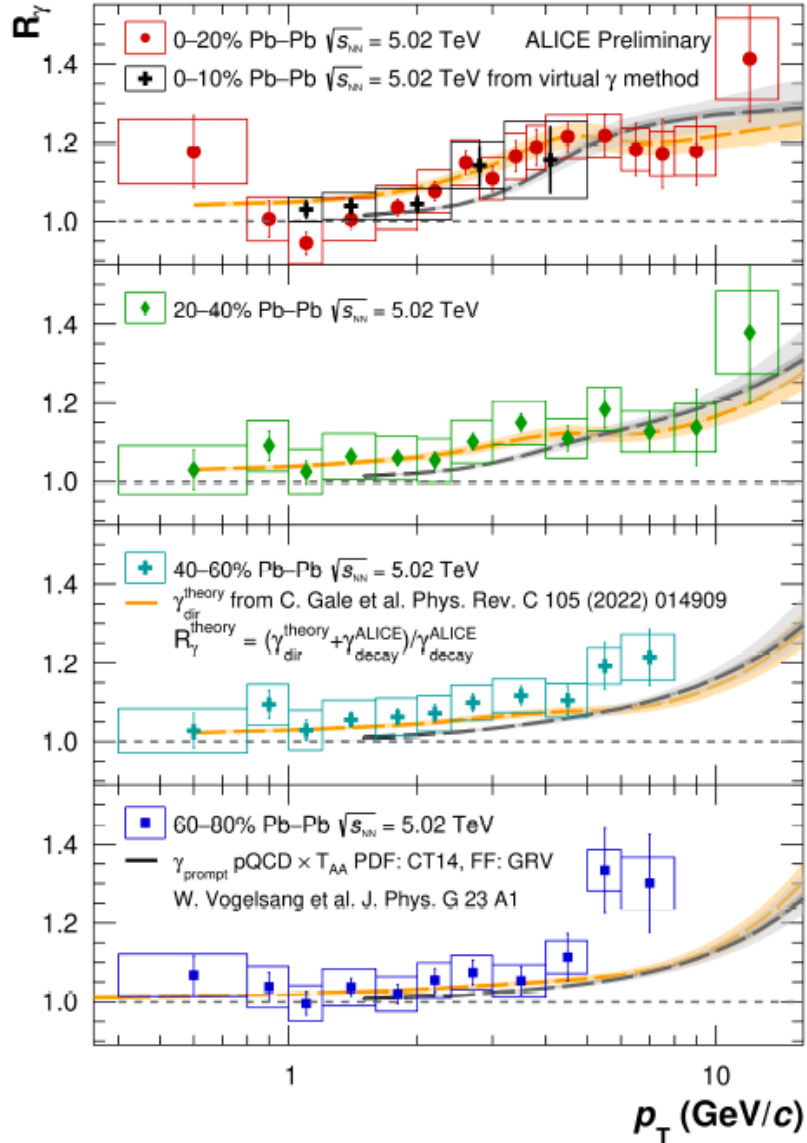
At low  $p_T$ :

$R_\gamma$  is close to 1  $\rightarrow$  small thermal contribution

For  $p_T > 2-3$  GeV/c:

- Excess which can be attributed to pre-equilibrium and prompt (hard scattering) photons

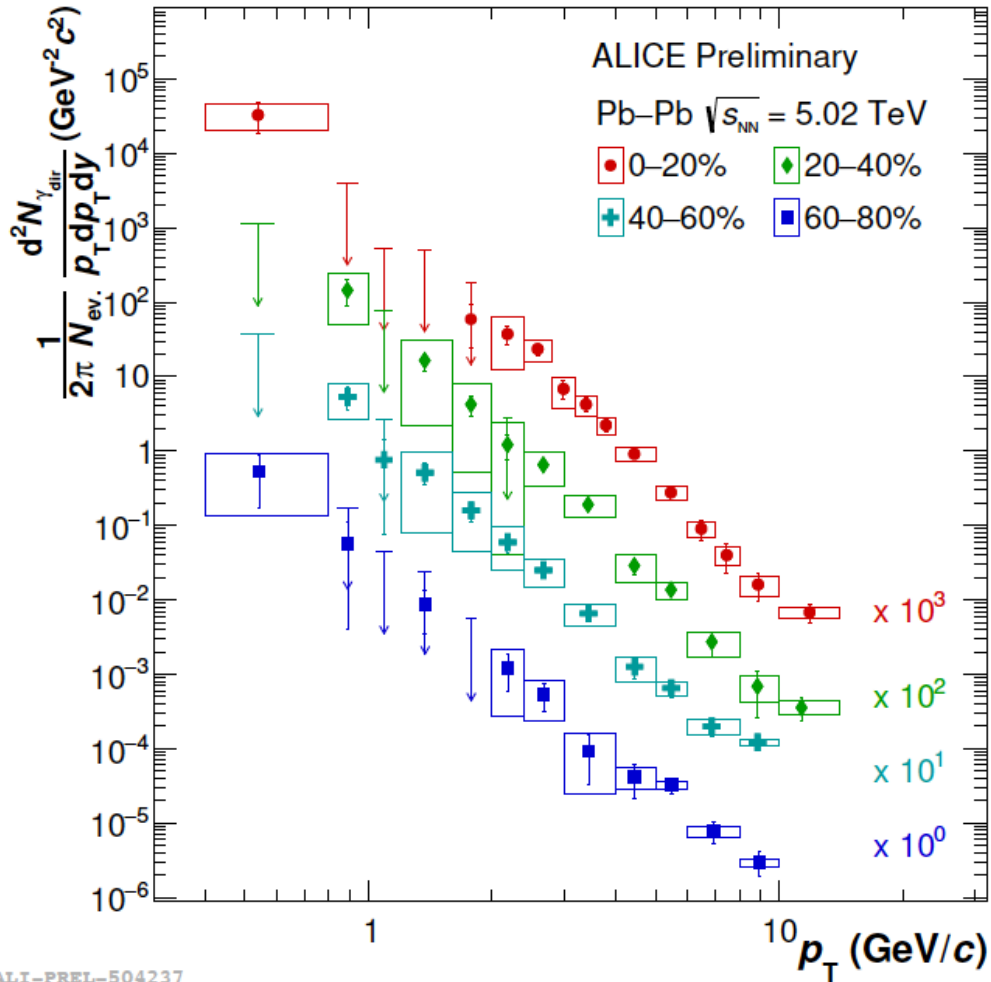
# QGP thermal emission: Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV



ALI-PREL-524126

ALI-PREL-514357

# Direct photon spectra in Pb-Pb at $\sqrt{s_{NN}} = 5.02$ TeV

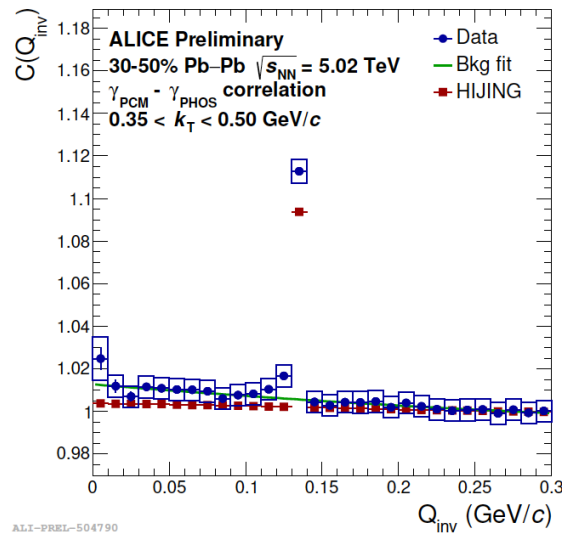
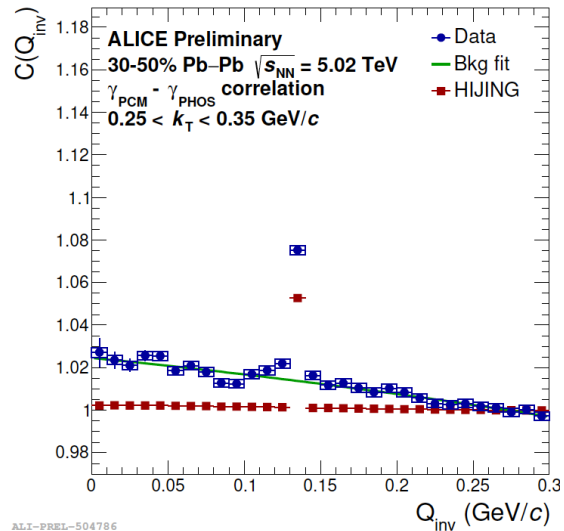
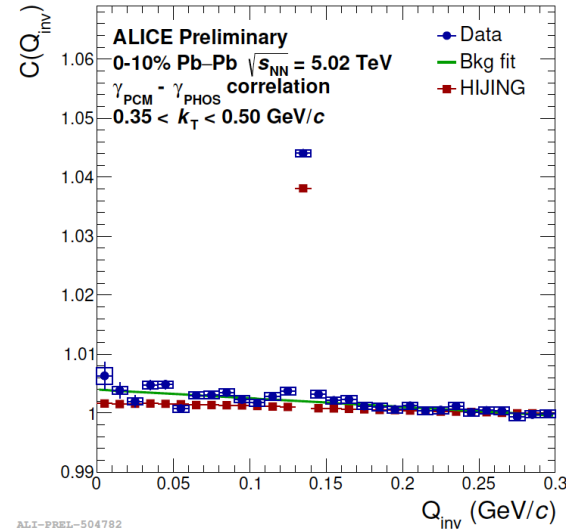
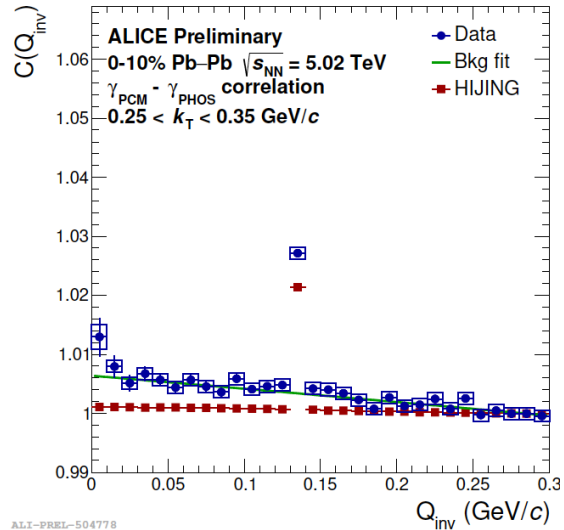


$$N_{\gamma,\text{dir}} = N_{\gamma,\text{inc}} - N_{\gamma,\text{dec}} = \left(1 - \frac{1}{R_\gamma}\right) \cdot N_{\gamma,\text{inc}}$$

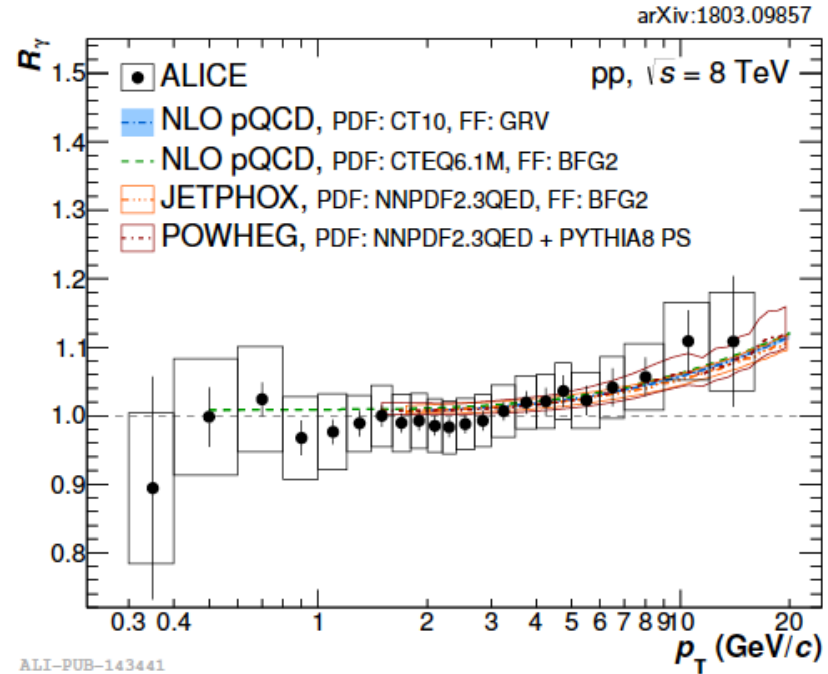
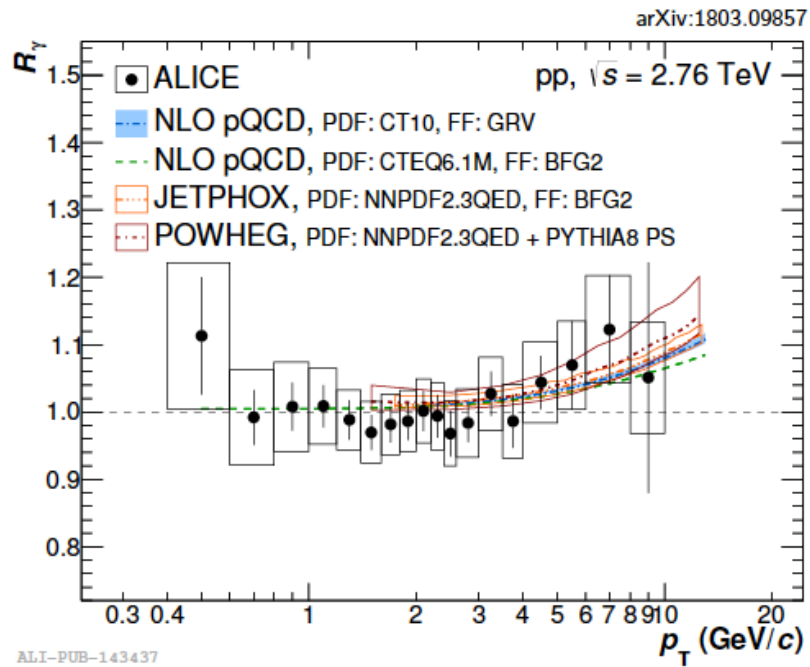
Upper limits (90% CL) given where  $\gamma_{\text{dir}}$  consistent with 0

ALI-PREL-504237

# Bose-Einstein $\gamma\gamma$ correlations in Pb-Pb collisions



# Direct $\gamma$ in pp collisions

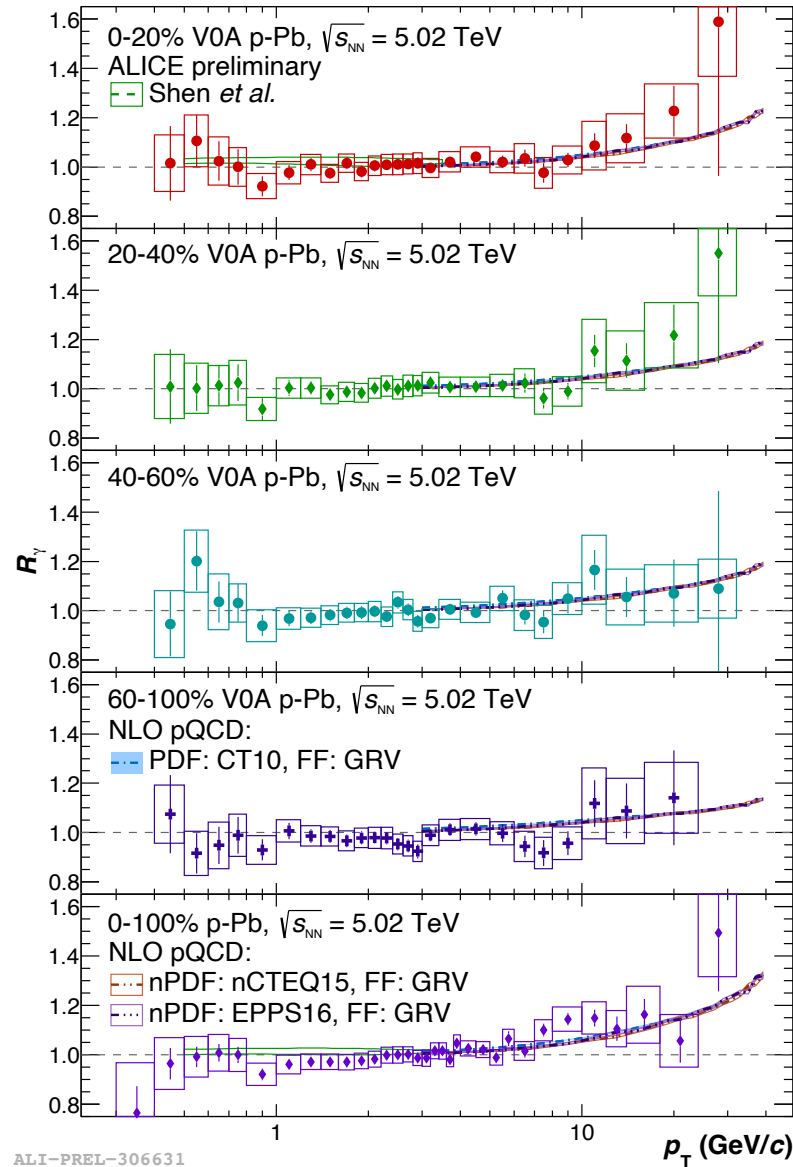


Combination of several reconstruction techniques via BLUE method.  
Theoretical NLO pQCD prediction plotted as

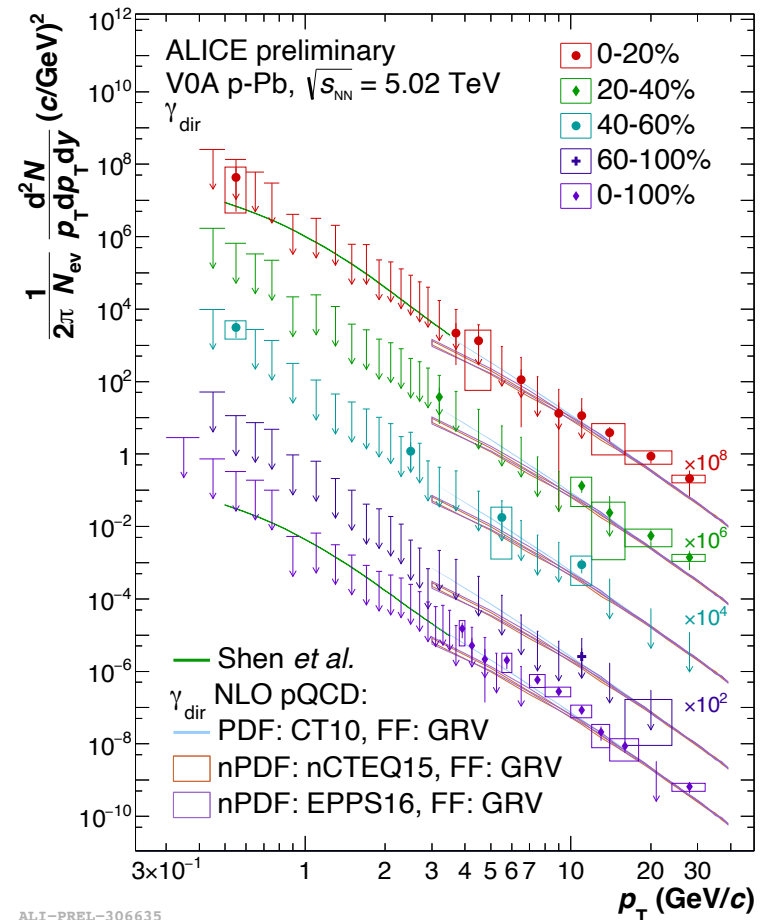
$$R_{\gamma}^{pQCD} = 1 + N_{coll} \frac{\gamma_{pQCD}}{\gamma_{decay}}$$

No significant excess observed at low  $p_T$ .  
About  $1 - 2\sigma$  deviation from unity for  $p_T > 7$  GeV/c

# Direct $\gamma$ in pPb collisions



ALI-PREL-306631



ALI-PREL-306635

No significant excess observed at low  $p_T$ .  
Accuracy is not yet sufficient to confirm/exclude thermal radiation in p-Pb collisions